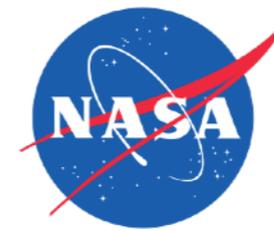
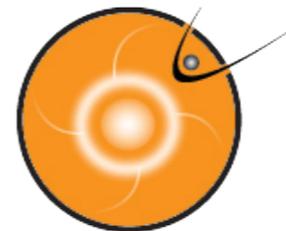


# Integrated Real-Time Modeling System for Heliospheric Space Weather Forecasting

*Dusan Odstrcil (PI), Lan Jian (Co-I), and Janet G. Luhmann (Co-I)*  
supported by:

*NASA/CCMC (L Mays, P. MacNeice, A. Taktakishvili)*

collaboration with: NOAA/SWPC & EU/HELcats



# Integrated Real-Time Modeling System for Heliospheric Space Weather Forecasting

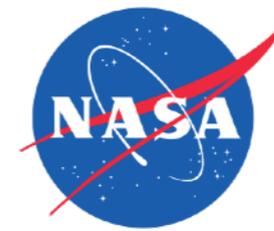
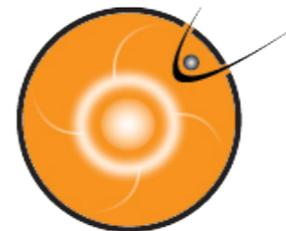
- Improvements in the WSA-ENLIL-Cone modeling system (*Dusan Odstrcil*)
- Verification and Validation of the updated WSA-ENLIL-Cone modeling system (*Lan Jian*)
- Modeling/predicting solar energetic particle (SEP) events using SEPMOD with ENLIL (*Janet Luhmann/Christina Lee*)
- Q & A

# Integrated Real-Time Modeling System for Heliospheric Space Weather Forecasting: Improvements in the WSA-ENLIL-Cone Modeling System

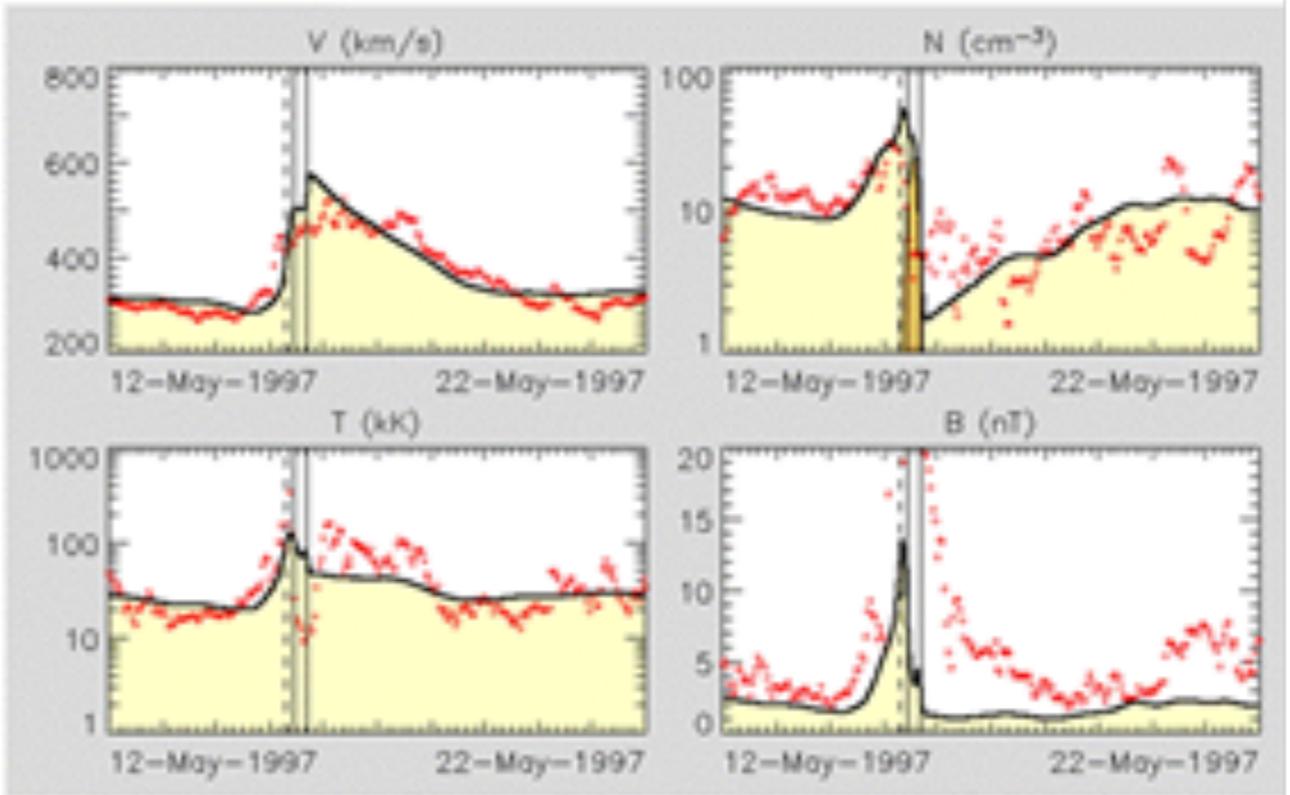
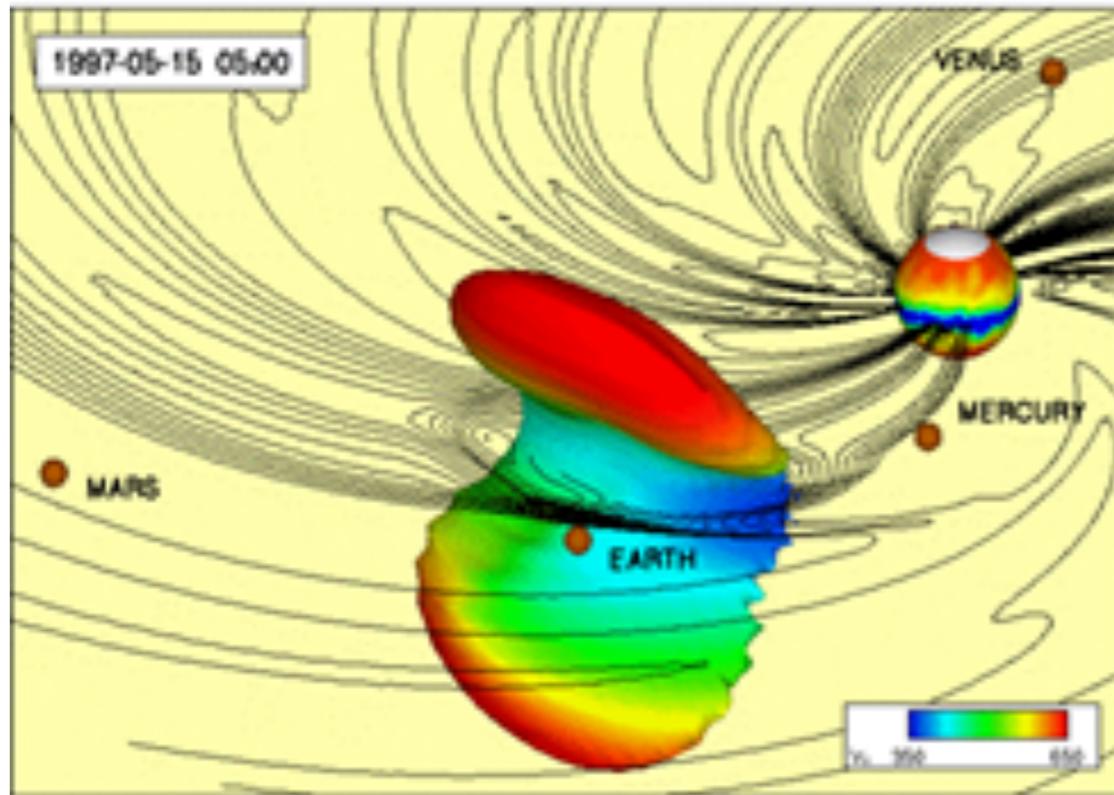
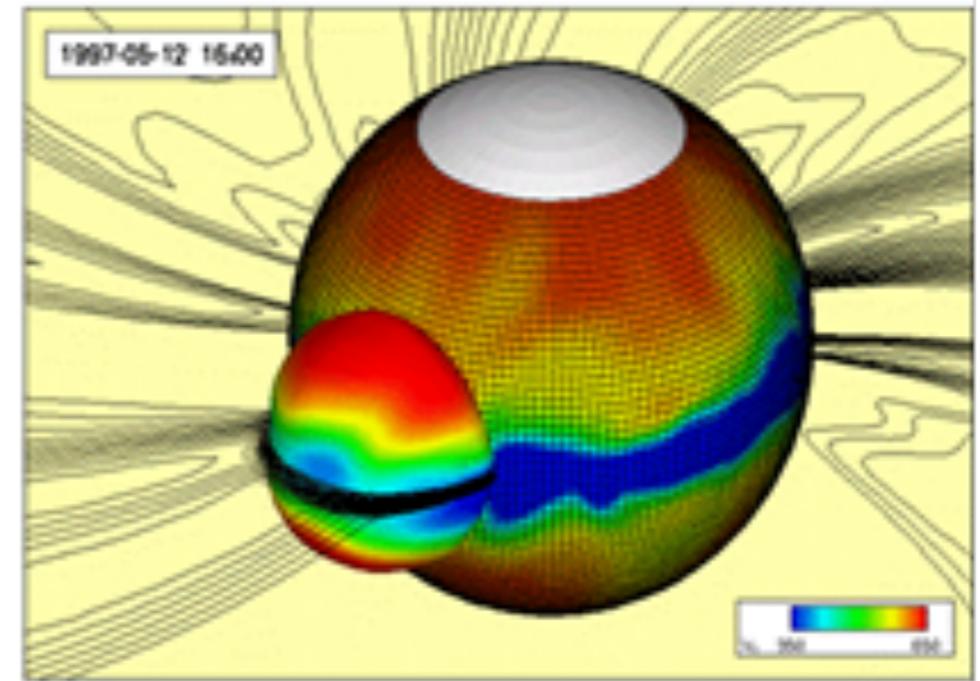
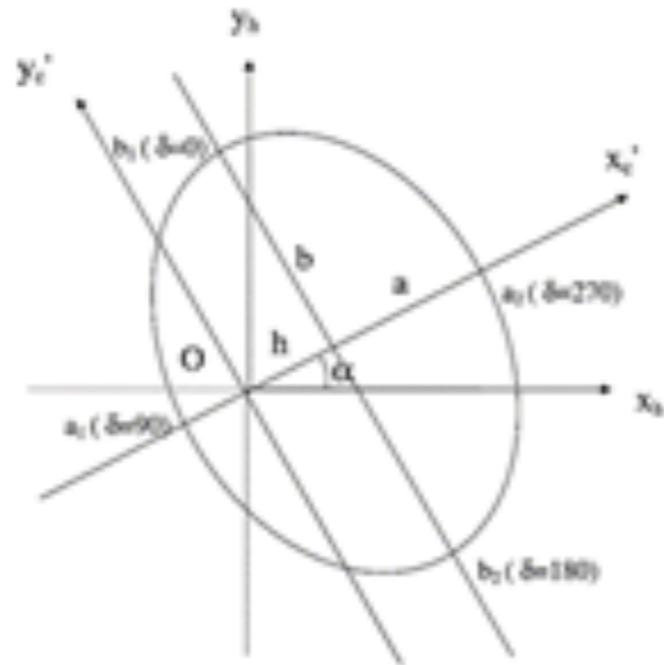
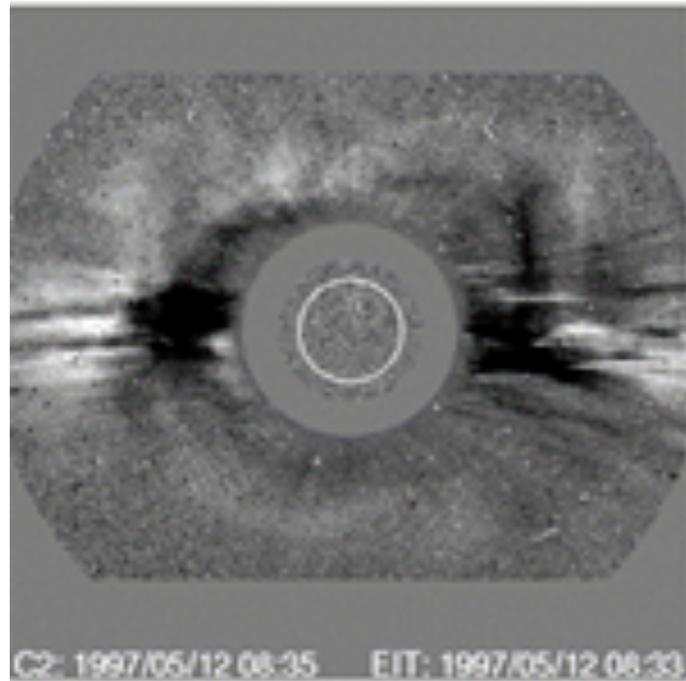
*Dusan Odstrcil (PI), Lan Jian (Co-I), and Janet G. Luhmann (Co-I)  
supported by:*

*NASA/CCMC (L Mays, P. MacNeice, A. Taktakishvili)*

*collaboration with: NOAA/SWPC & EU/HELICATS*

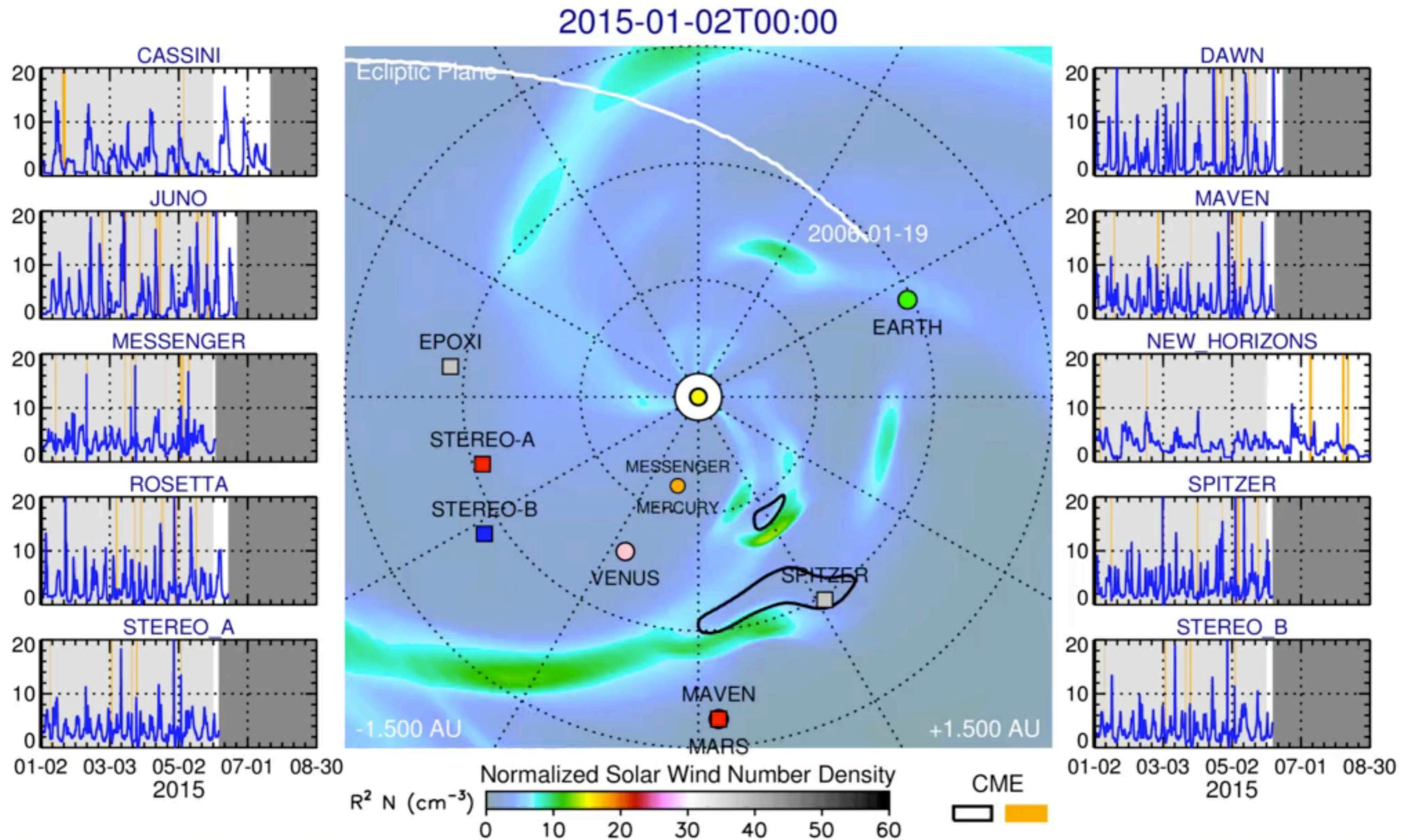


# WSA-ENLIL-Cone Modeling System — Heliospheric Space Weather



- Observationally driven, near-real time, “hybrid” modeling system for heliospheric space weather
- Routine simulation of co-rotating streams & CMEs, event-by-event, much faster than real-time
- Supported by NSF.CISN and implemented at NASA/CCMC & NOAA/SWPC

# WSA-ENLIL-Cone — Heliospheric Predictions for NASA Missions



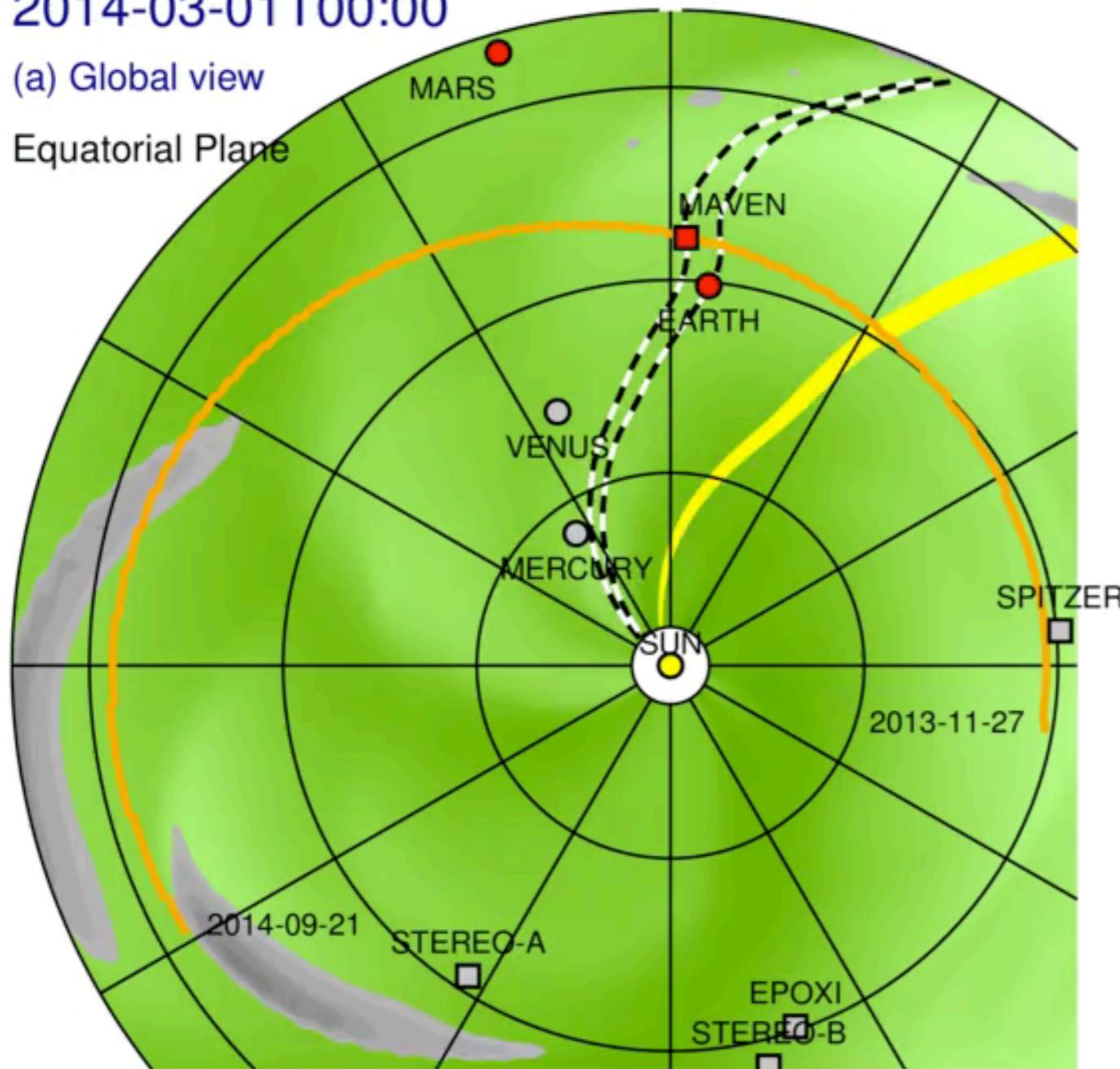
- All CMEs ( $>500$  km/s) fitted by CCMC in past 8 months are used for 4-months prediction at NH
- History (light-grey background) and prediction (white background) for heliospheric missions
- Can be used for mission planning and operational support at NASA/CCMC

# MAVEN Cruise to Mars — Predicted Shock-SEP Alerts

2014-03-01T00:00

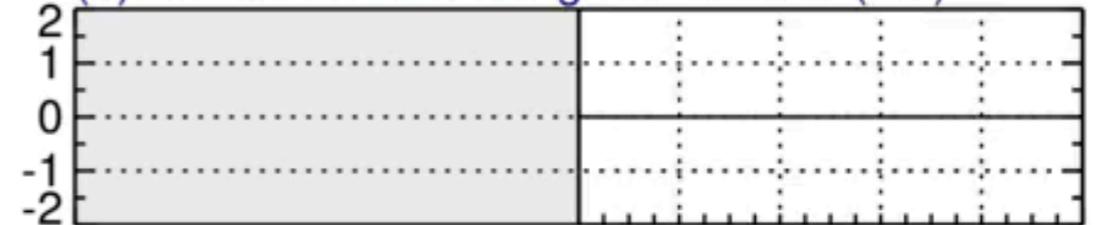
(a) Global view

Equatorial Plane

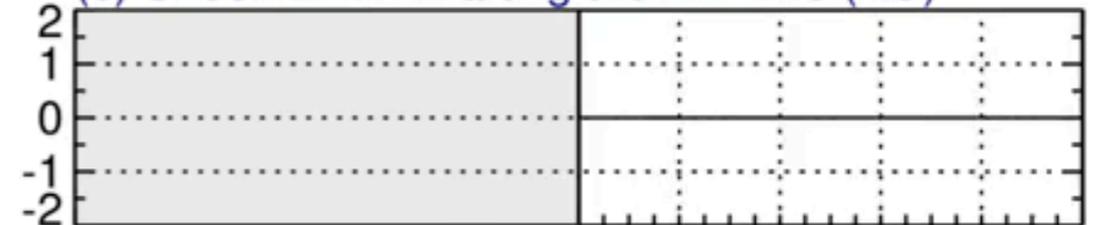


2014-03-01T00 + 0.00 days

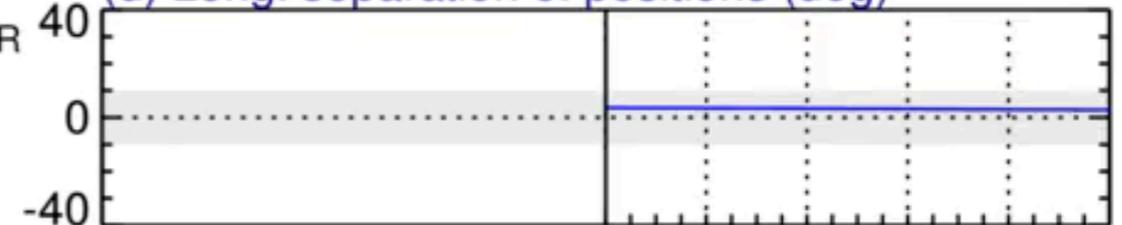
(b) Shock-MAVEN along the IMF line (AU)



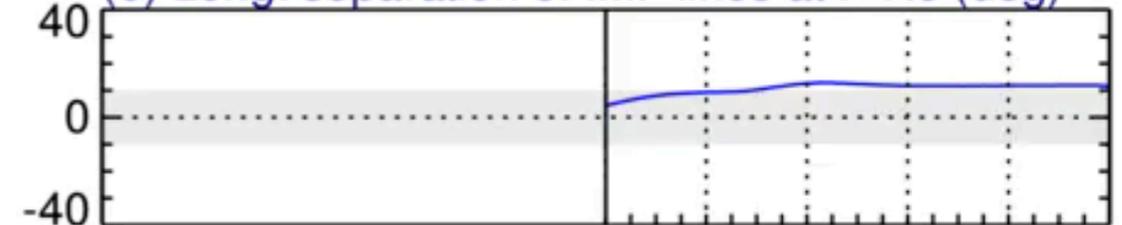
(c) Shock-EARTH along the IMF line (AU)



(d) Long. separation of positions (deg)



(e) Long. separation of IMF lines at  $r=R_e$  (deg)



01 02 03 04 05 06  
2014-03

Vamb (km/s) 0 400 800

N/Namb 1.5 4.0 6.5

V-Vamb at shock (km/s) 200 400 800

IMF line  
- - -

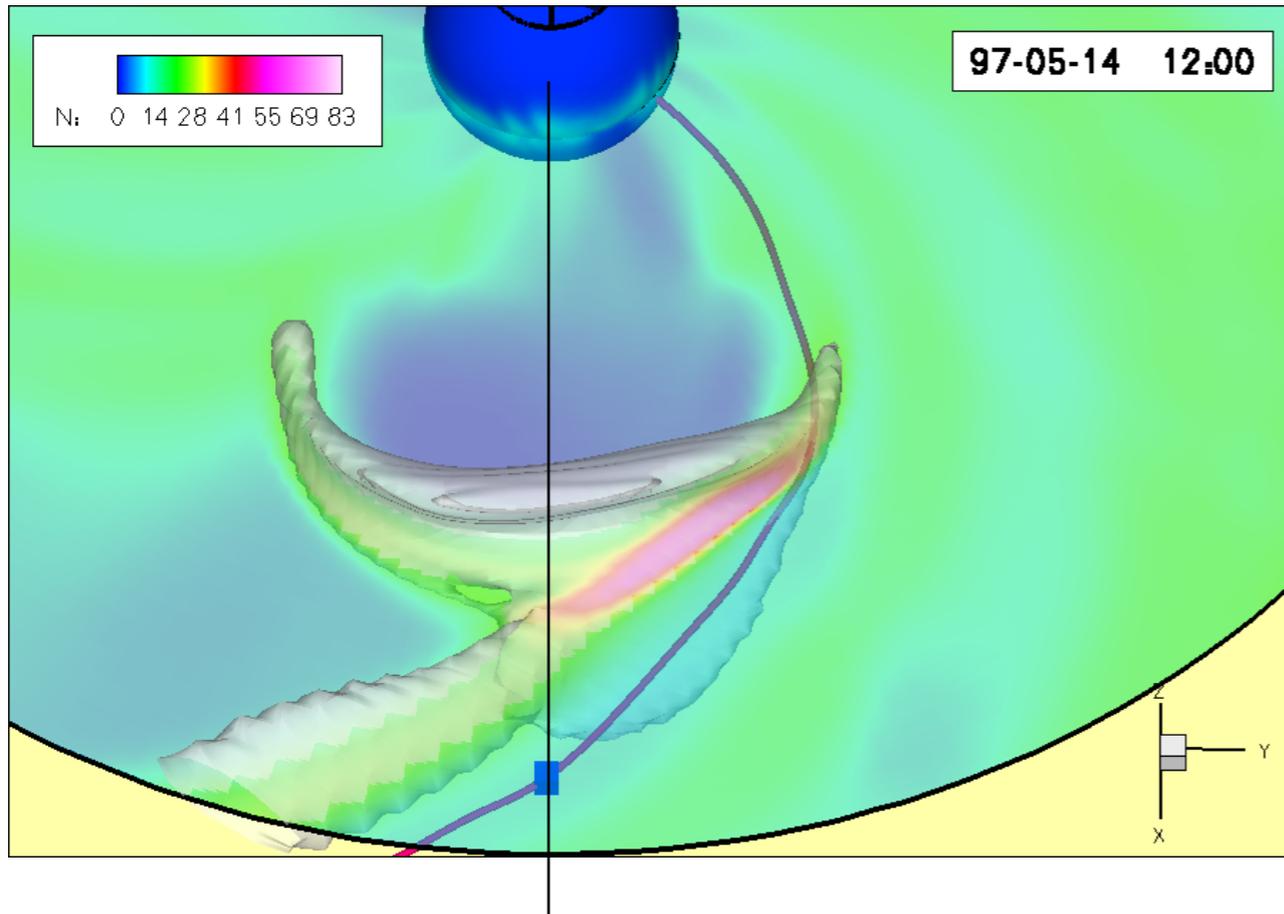
ENLIL-lowres + WSADT-GONGb + Cone-SWRC

HELIO WEATHER

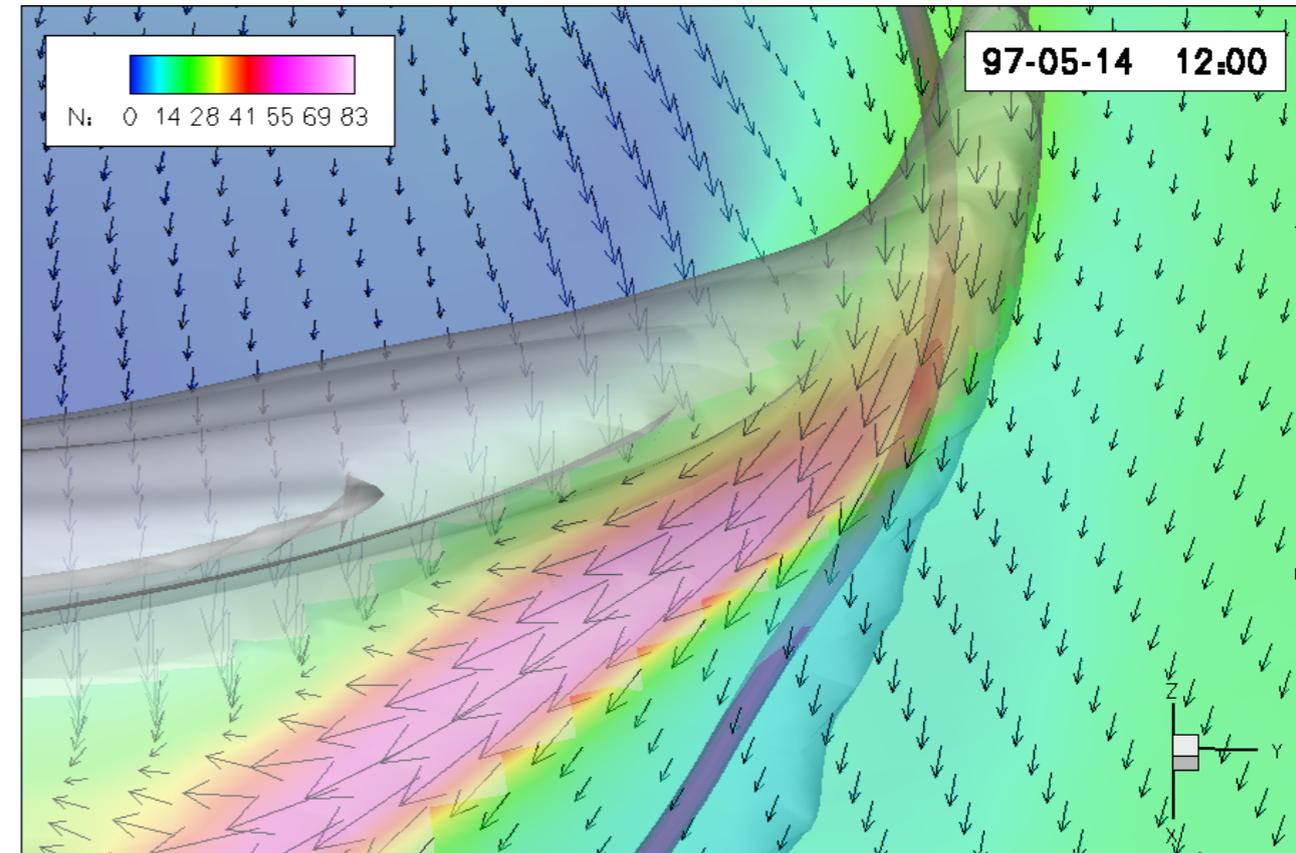
- All (“classical-propulsion”) missions to Mars follow the Hohmann trajectory
- Spacecraft close to IMF line passing through Earth with SEP measurements for alerts
- Simulations confirm Posner’s idea (PSS, 2013) except periods when IMF is disturbed by CMEs

# Shock Detection Challenge

Global view

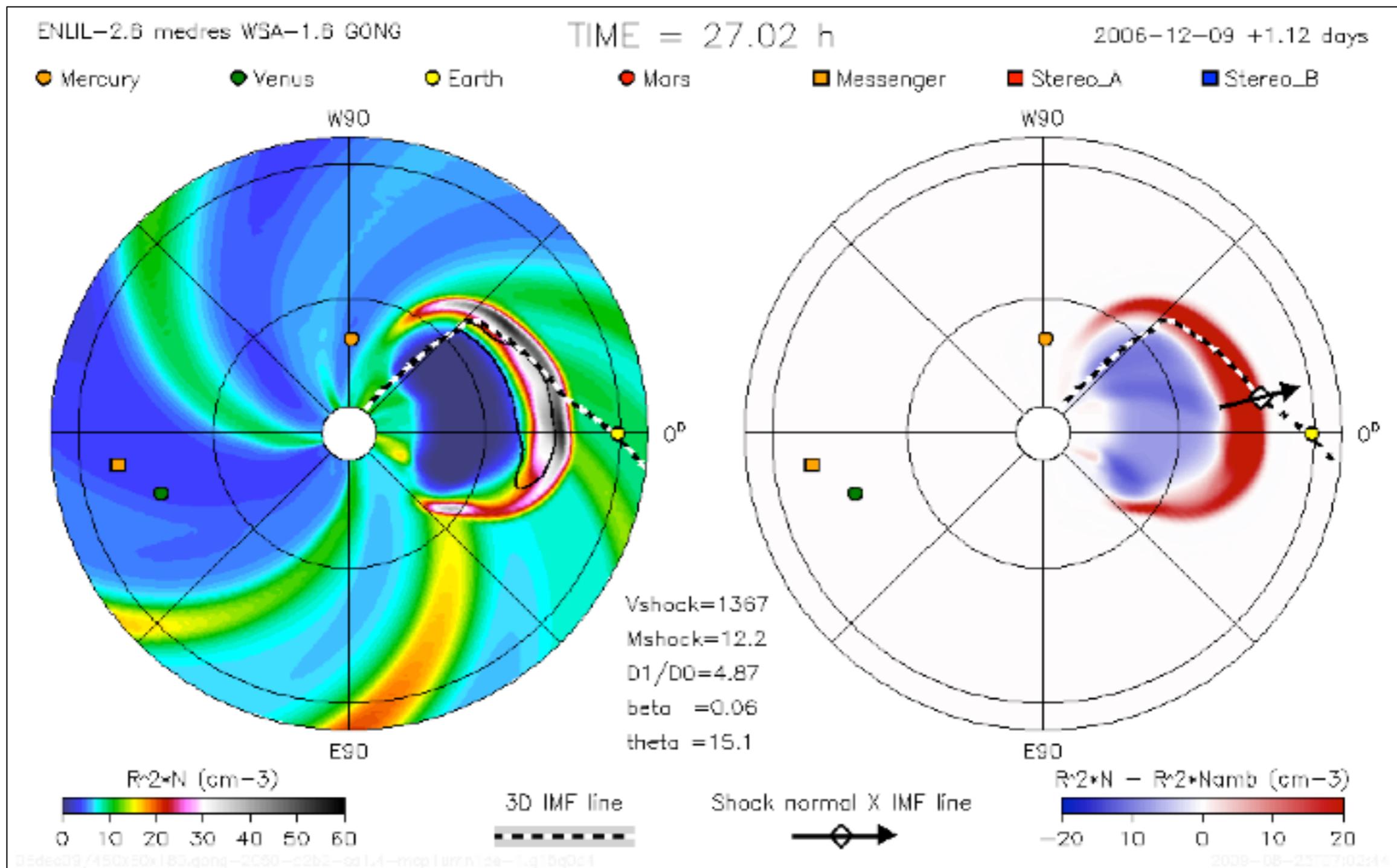


Detailed view



- IMF line connects geospace with an interplanetary shock under very large inclination angle because of: (1) spiraling IMF line and (2) bow-shaped shock front
- Thus determination of shock parameters from MHD values stored along the IMF line is very difficult because many numerical grid points are used across the shock structure and pre- and post-shock values are at different solar wind

# Automatic Detection of CME-Driven Shocks



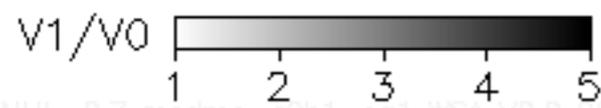
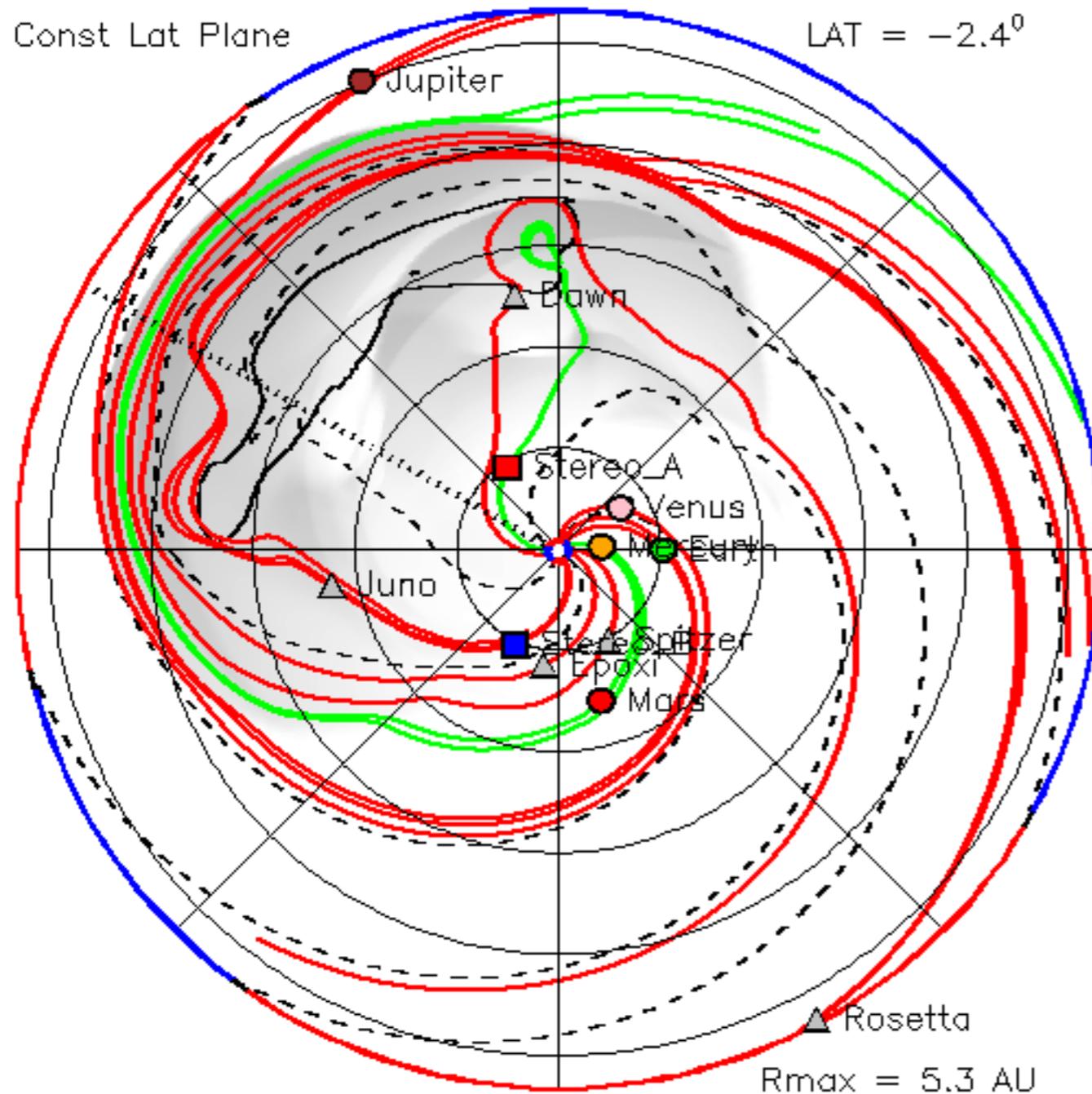
- Two simulations are used: one for background only and the other for background+transient
- Subtraction the two results to clearly identify an interplanetary shock leading edge along the observer connected IMF line

# WSA-ENLIL-Cone Modeling System — Project Proposal

2012-07-30T06:00

Const Lat Plane

LAT =  $-2.4^\circ$



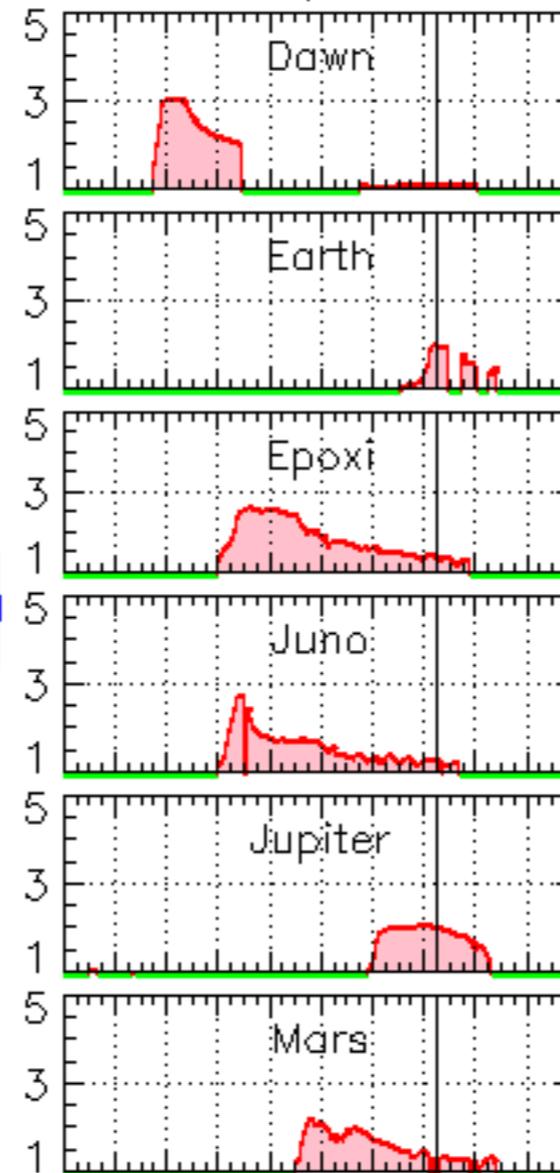
IMF polarity



Current sheath



V1/V0



23 25 27 29 31 02  
2012-07/2012-08

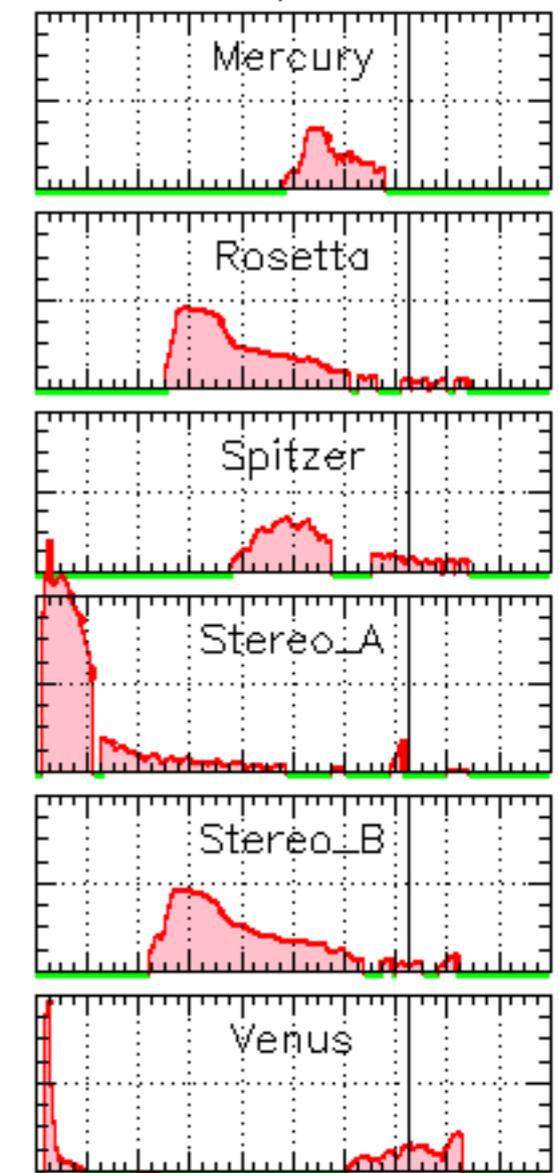
IMF line



IMF+SEP line



V1/V0



23 25 27 29 31 02  
2012-07/2012-08

ICME direct



ICME ejecta



ENLIL-2.7 medres-69b1-s01 WSA\_V2.2 GONG-2126\_306

2012-07-23/nun.1760x60x180x2.a9b1-s01.22-mcp1van2cd-1.a53q5.gong-2012-07-23T00 2012-08-17

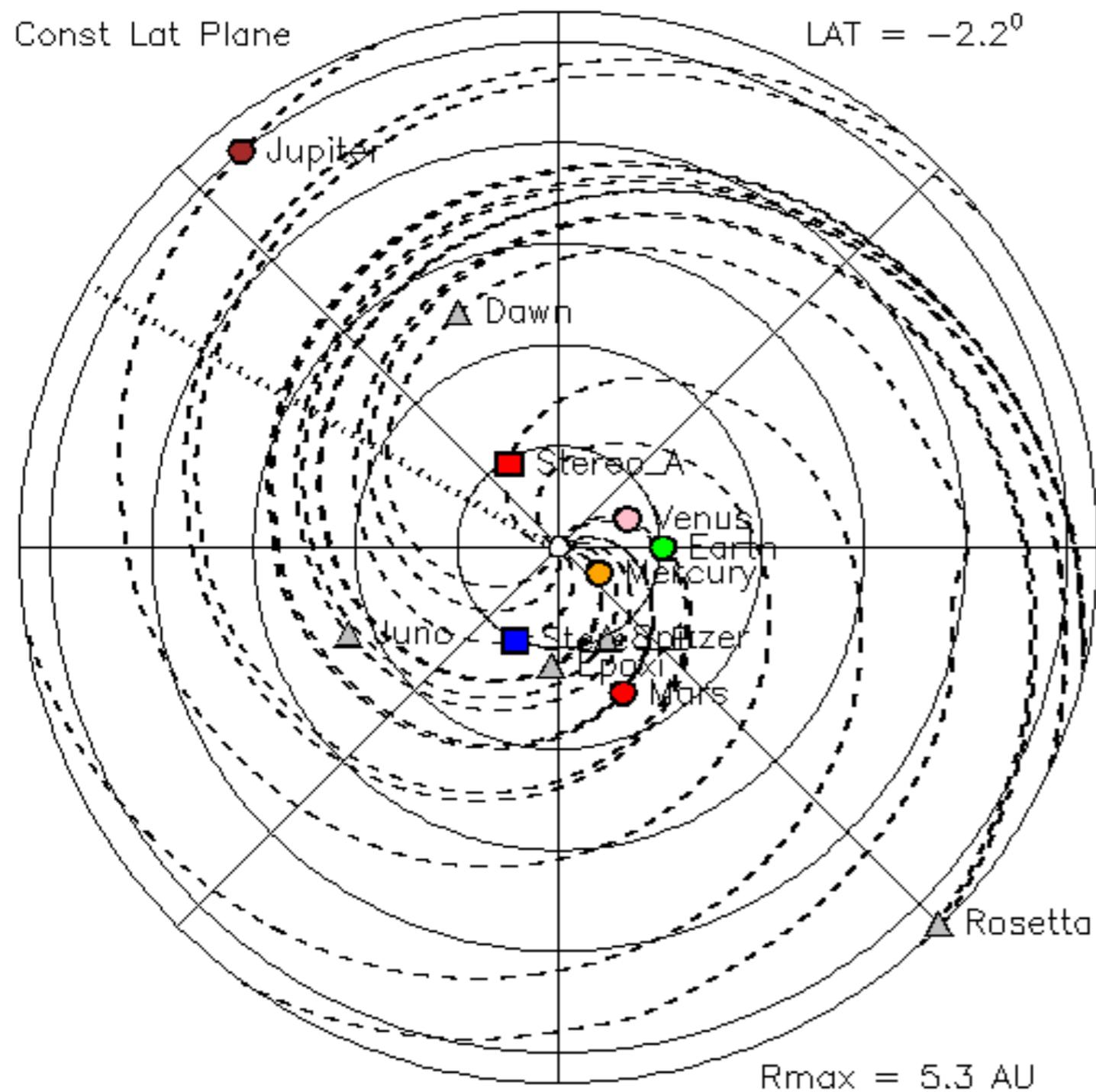
- Driven by evolving WSA maps & improved CME fittings — predictions in mid heliosphere
- Development of additional products and applications (SEP, white-light, etc.)
- Validation across the solar-activity cycle

# Multi-CME Events — 2012 July

2012-07-12T00:00

Const Lat Plane

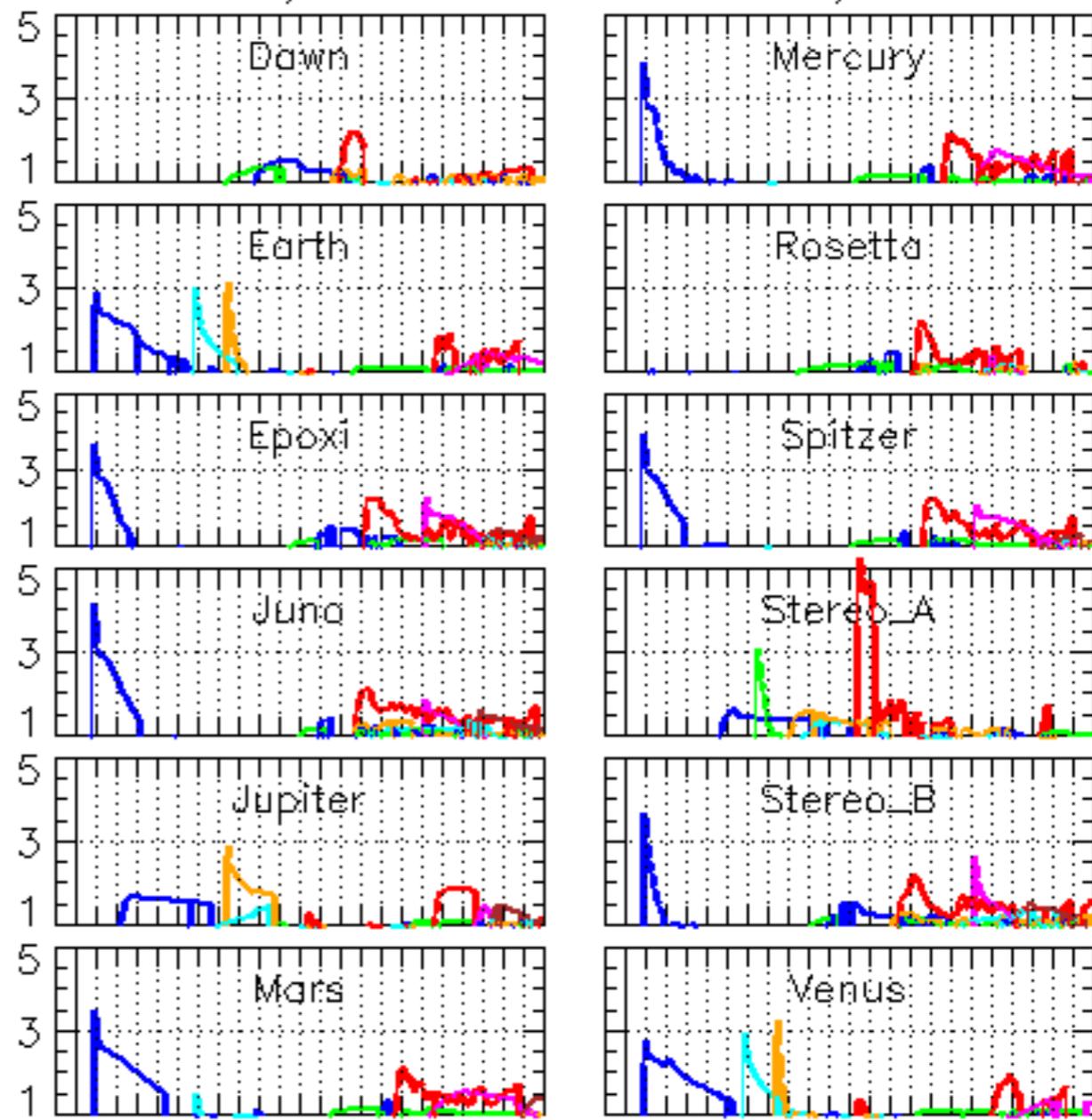
LAT =  $-2.2^\circ$



2012-07-11T00 - 11.00 days

V1/V0

V1/V0



11 14 17 20 23 26 29 01 04 11 14 17 20 23 26 29 01 04  
2012-07/2012-08 2012-07/2012-08

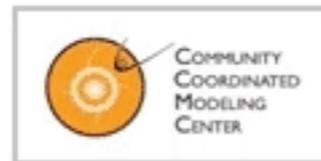
IMF line IMF+SEP line ICME direct ICME ejecta



# WSA-ENLIL-DONKI-HELCASTS — Solar Wind and CMEs

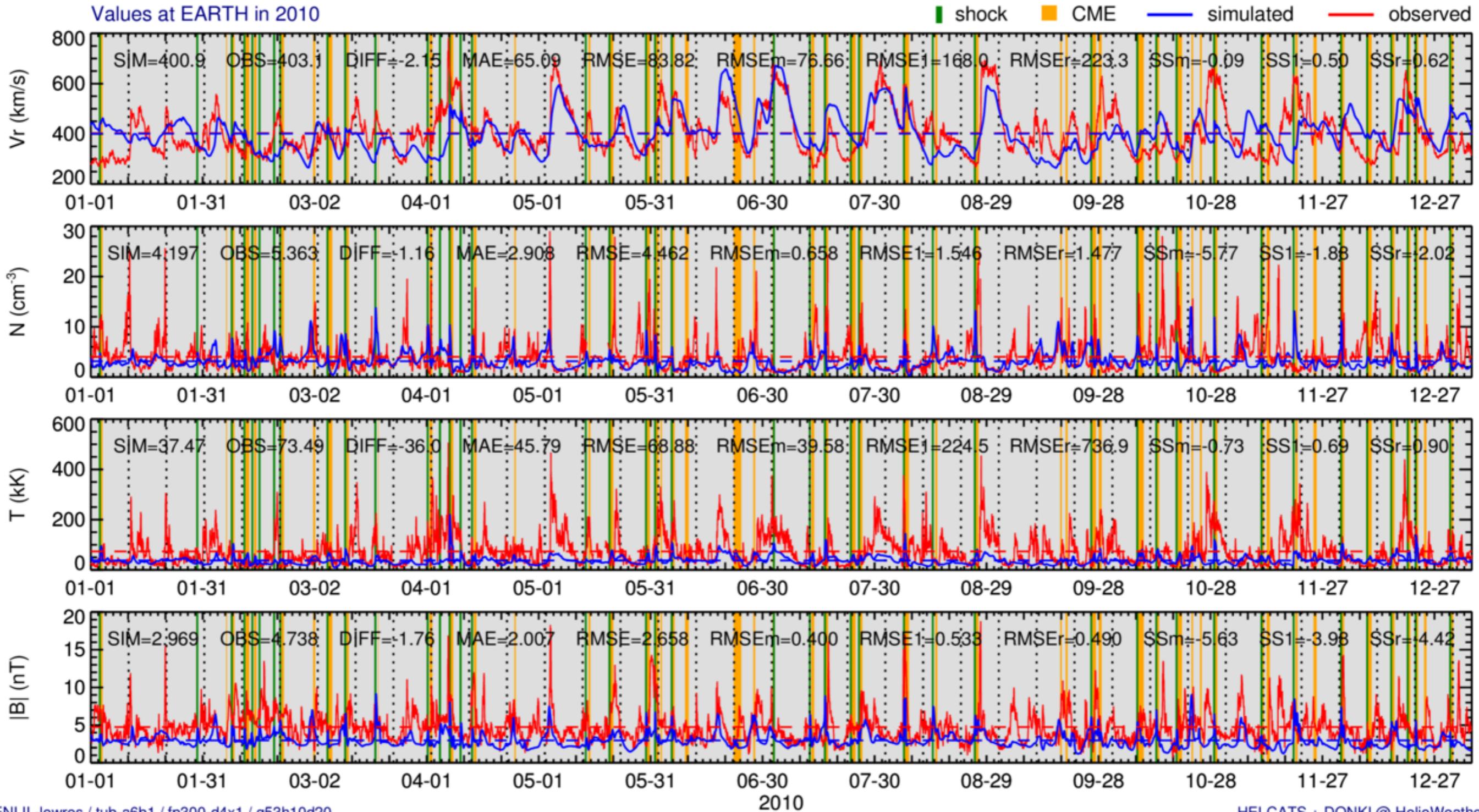
<http://helioweather.net/archive>

Year	Month											
<a href="#">2007</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>
<a href="#">2008</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>
<a href="#">2009</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>
<a href="#">2010</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>
<a href="#">2011</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>
<a href="#">2012</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>
<a href="#">2013</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>
<a href="#">2014</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>
<a href="#">2015</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>
<a href="#">2016</a>	<a href="#">Jan</a>	<a href="#">Feb</a>	<a href="#">Mar</a>	<a href="#">Apr</a>	<a href="#">May</a>	<a href="#">Jun</a>	<a href="#">Jul</a>	<a href="#">Aug</a>	<a href="#">Sep</a>	<a href="#">Oct</a>	<a href="#">Nov</a>	<a href="#">Dec</a>

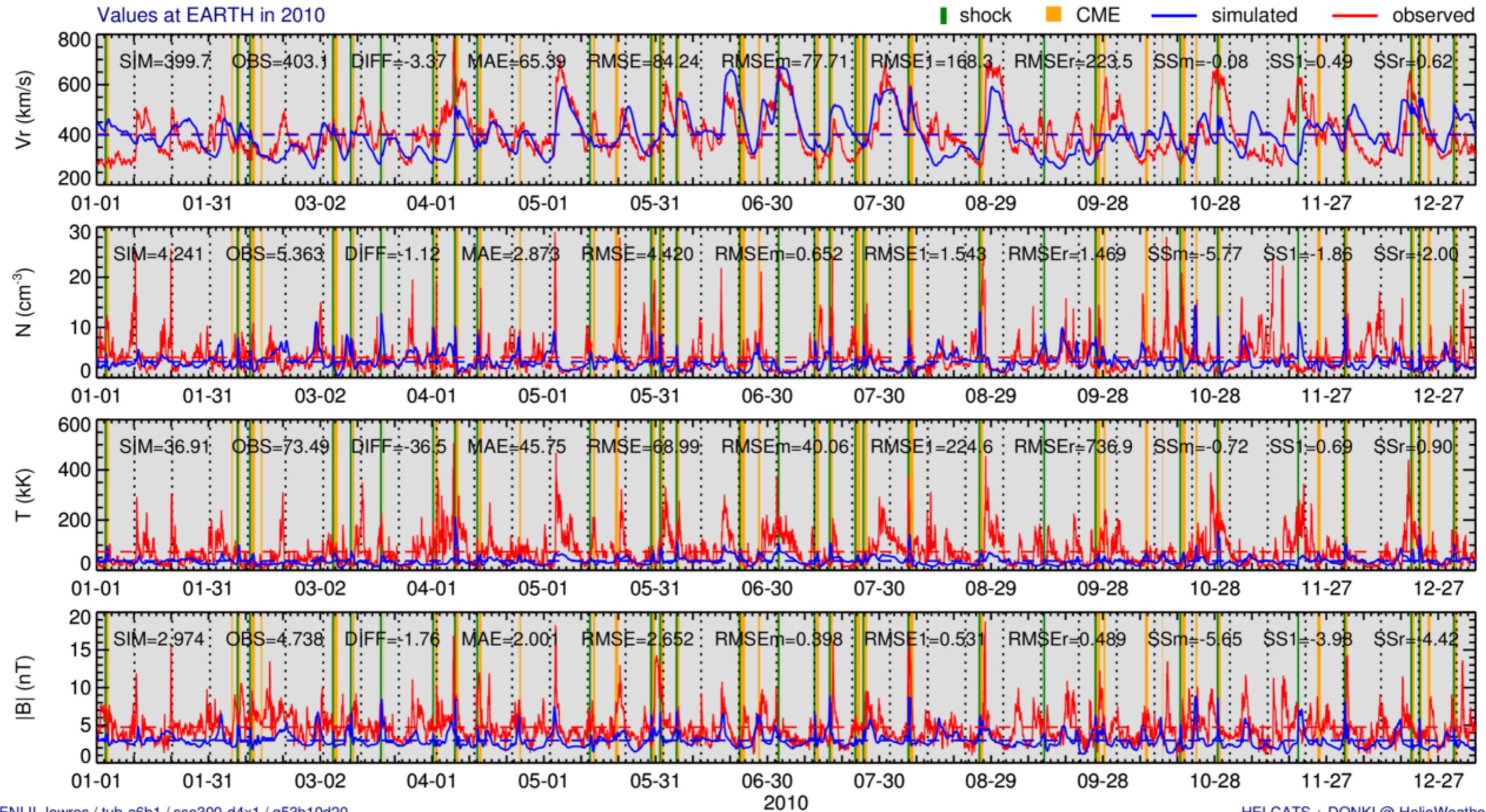


- Simulations of all CMEs from 2007 (STEREO data available) until present:
  - evolving background using the WSA maps
  - hydrodynamic ejecta by DONKI and HELCASTS fitting of CMEs > 300 km/s
- This covers basically the solar-activity cycle:
  - verification & validation during minimum and maximum solar activity
  - calibration of model free parameters

# WSA-ENLIL-DONKI-HELGCATS — CMEs > 300 km/s by FP in 2010



# WSA-ENLIL-DONKI-HELGCATS — CMEs > 300 km/s by SSE in 2010



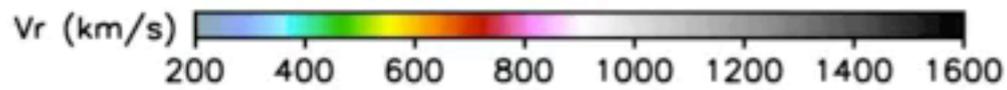
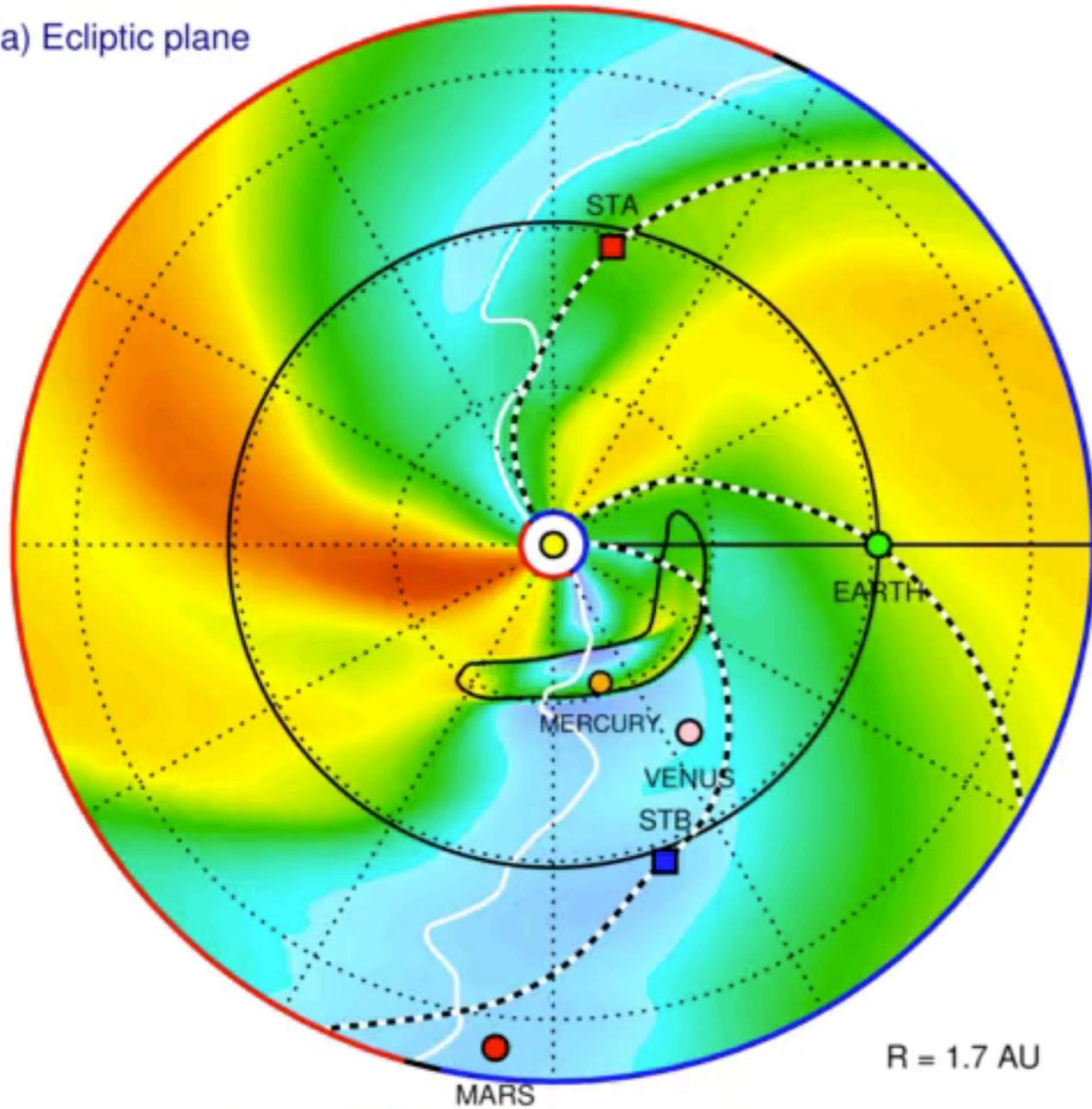
# WSA-ENLIL-DONKI-HELICATS — Solar Wind Speed in 2012-03

2010-08-01T00:00

EARTH

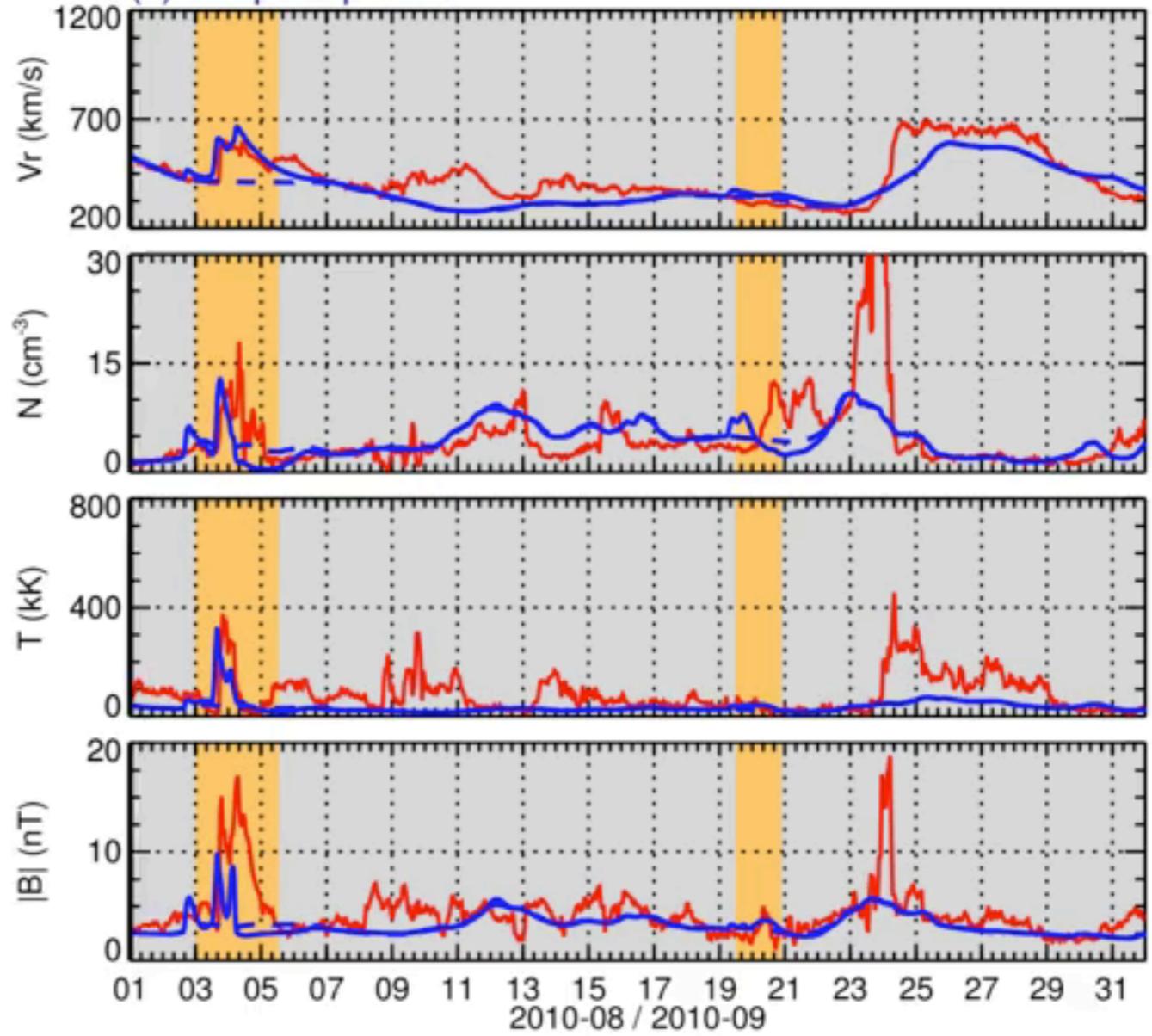
2010-08-01T00 + 0.000 days

(a) Ecliptic plane



CME

(b) Temporal profiles



IMF line

IMF polarity

HCS

CME

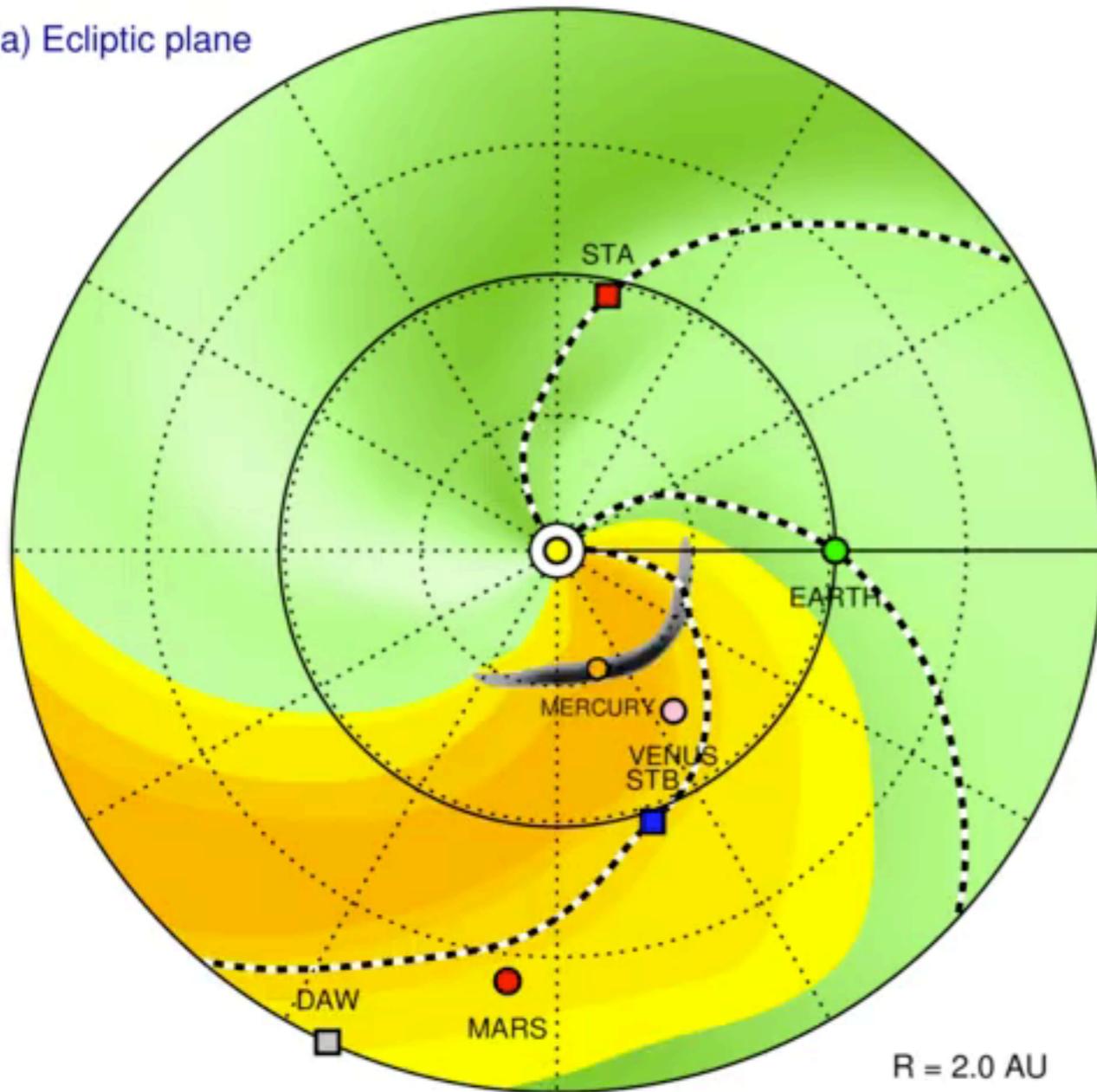
measured

simulated

# WSA-ENLIL-DONKI-HELICATS — Shock SEP Alerts

2010-08-01T00:00

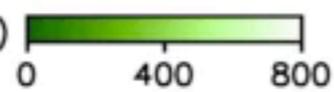
(a) Ecliptic plane



IMF line



Vomb (km/s)



N/Nomb



V-Vomb at shock (km/s)

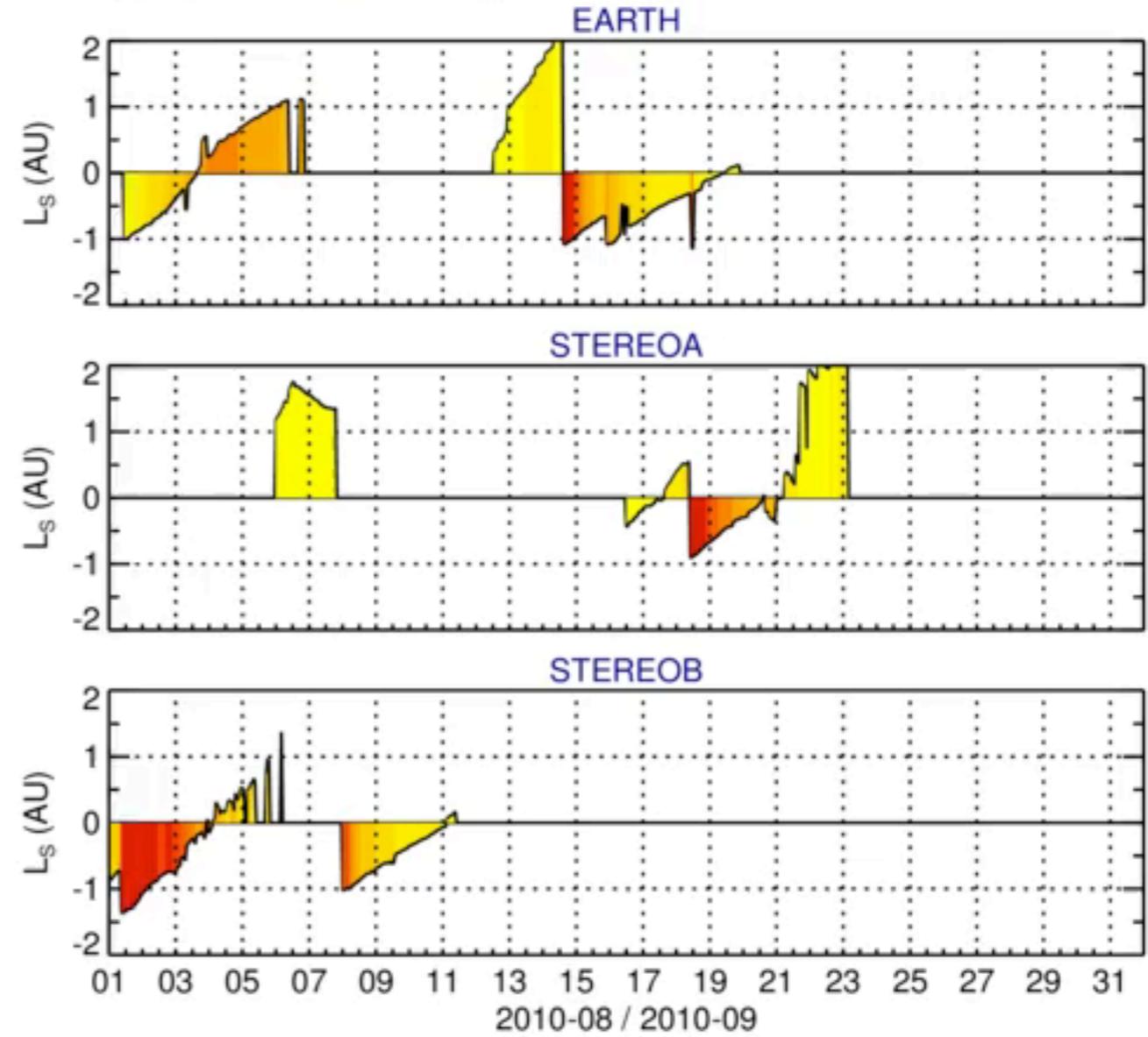


ENLIL-lowres + GONGb-WSAtu / fp400-d4x1 / g53h10

HELICATS + DONKI @ HelioWeather

2010-08-01T00 + 0.000 days

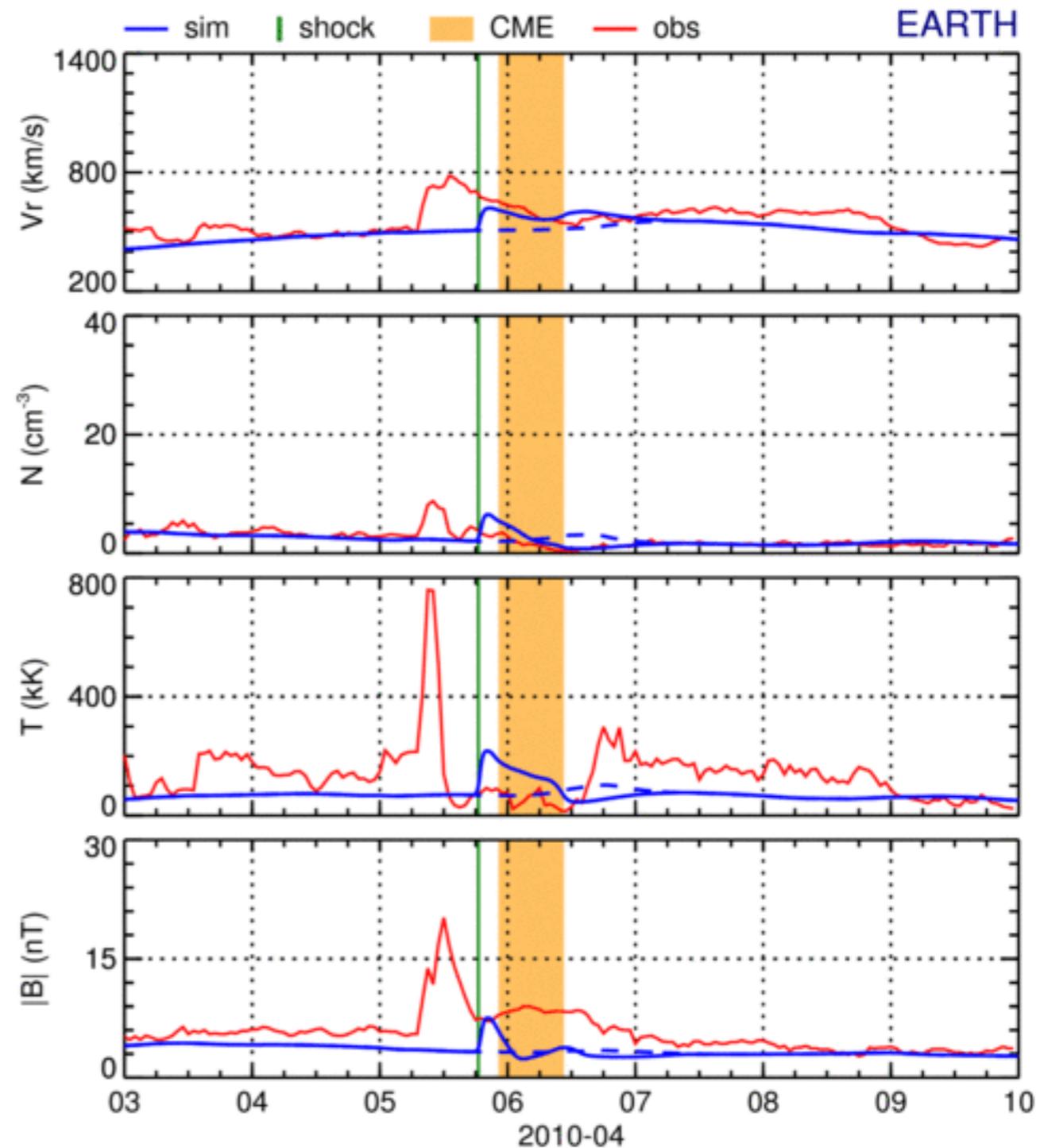
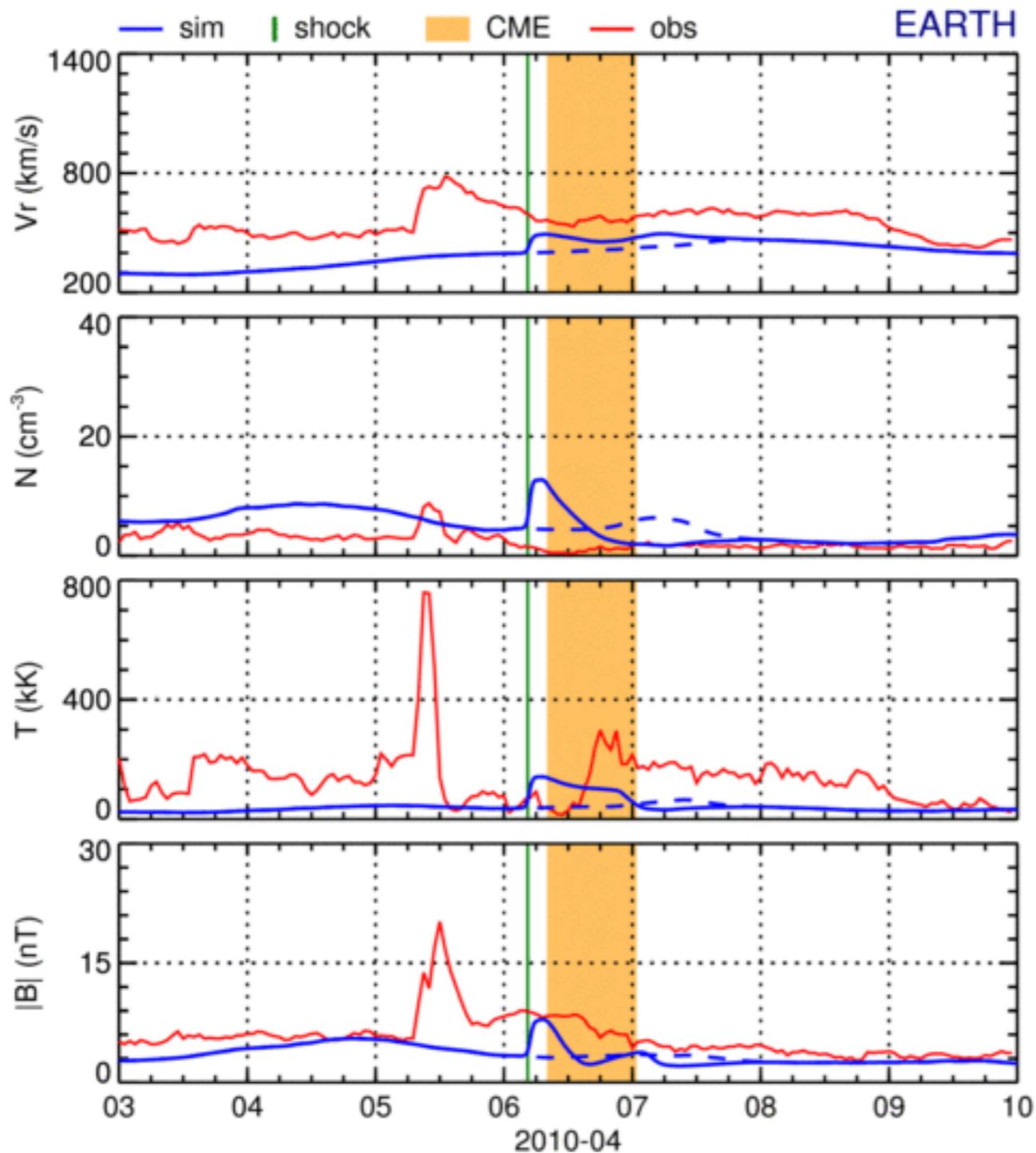
(b) Shock distance along the IMF line



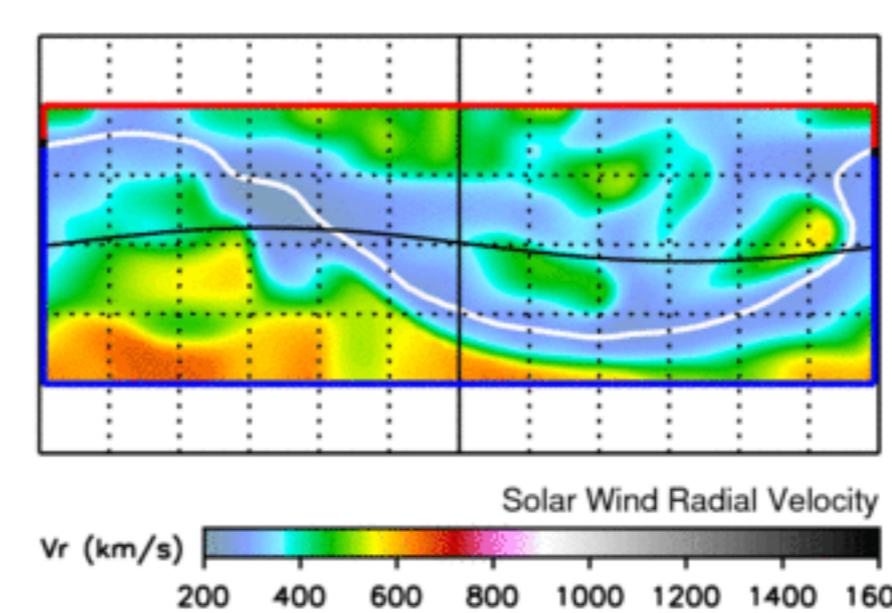
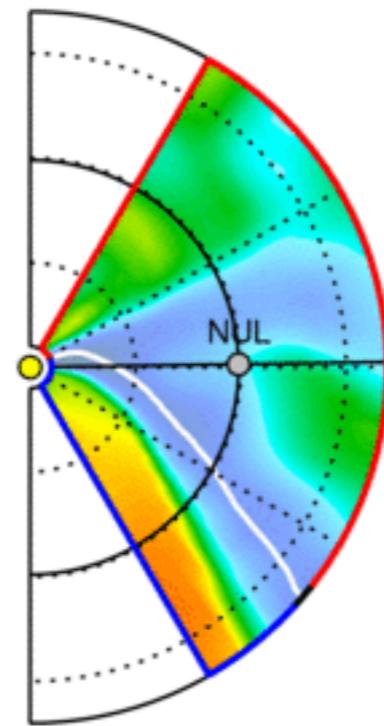
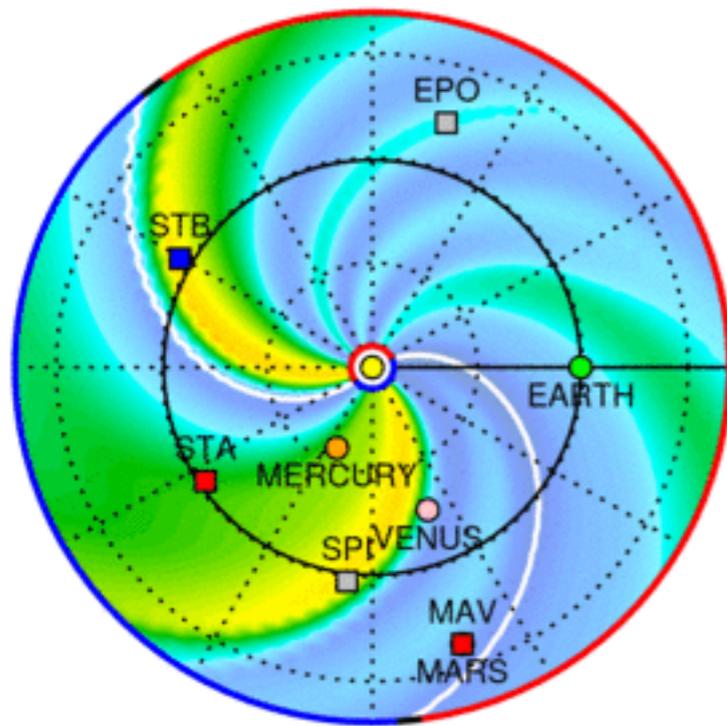
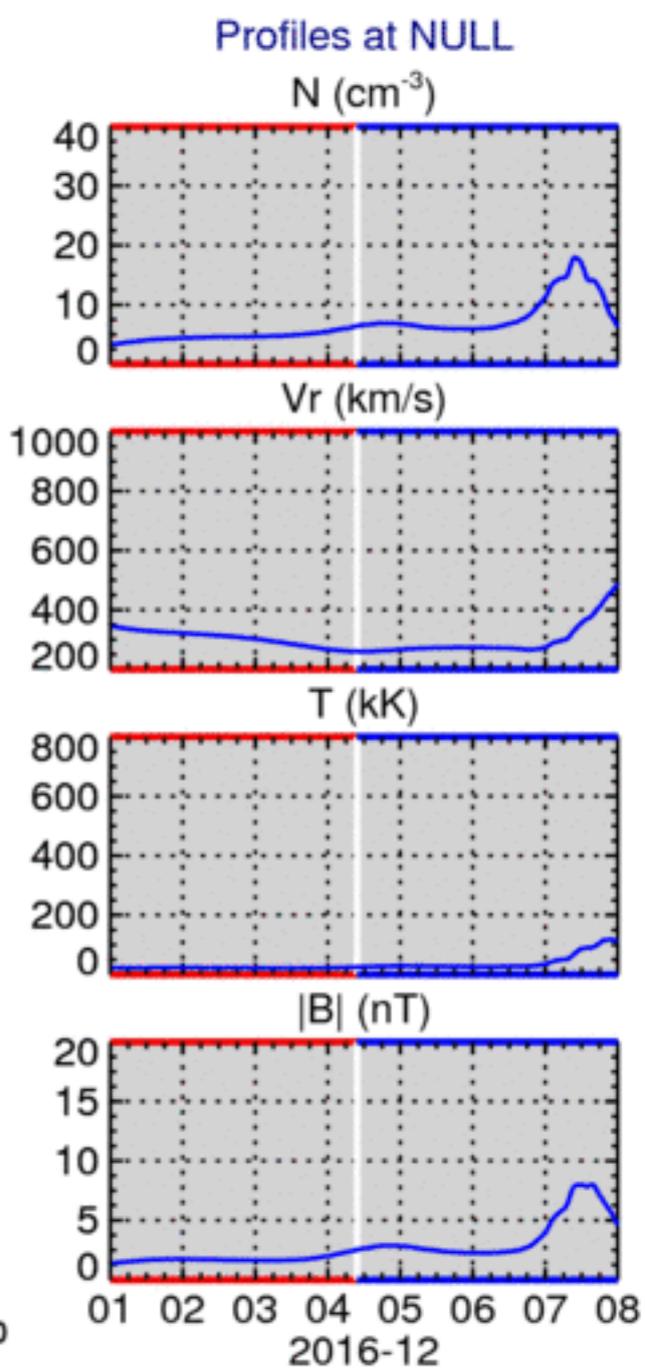
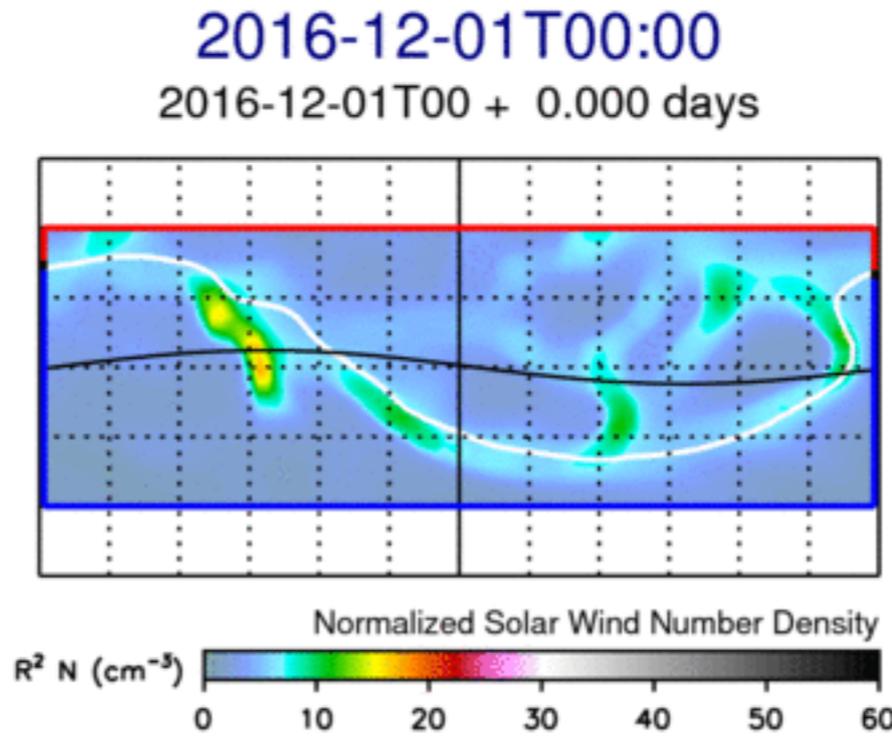
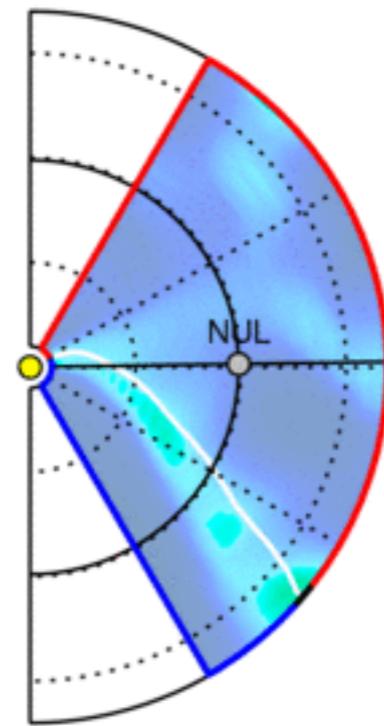
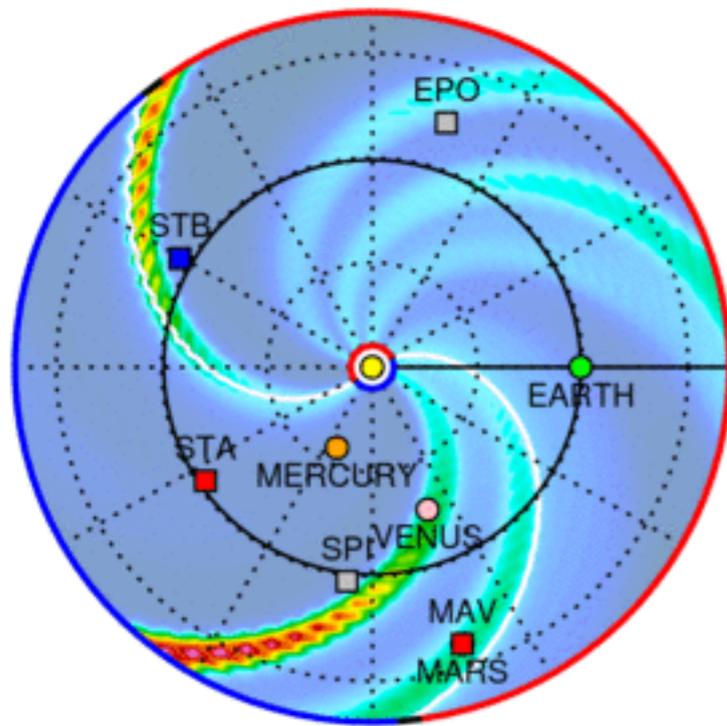
# iCCMC Event "A" — CME 2010-04-03

observed arrival: 2010-04-05T08:26, predicted arrival: 2010-04-06T03:30 (19.07 h later )

- “ad hoc” solar wind — predicted arrival: 2010-04-06T04:27 (20.02 h later)
- “realistic” solar wind — predicted arrival: 2010-04-05T18:32 (10.09 h later)



# 12 WSADA Maps — Solar Wind Speed on 2016-12-01

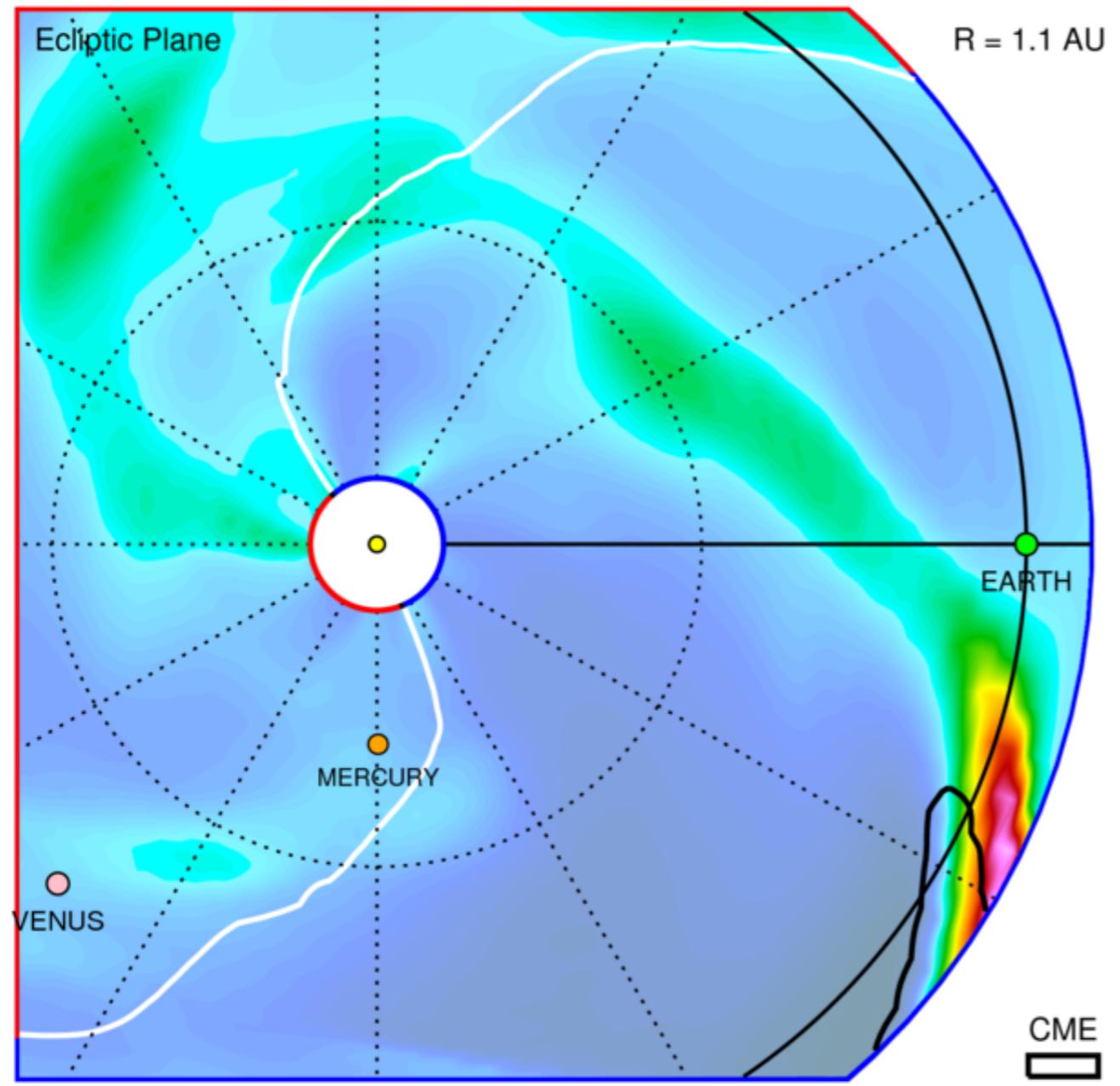


IMF polarity  
- [blue/red] +  
HCS [white line]

simulated  
[blue line]  
HeliWeather  
[red line]

# WSA-ENLIL-DONKI-HELCASTS — Solar Wind Speed in 2012-03

2010-04-03T00



$r^2 N$  ( $\text{cm}^{-3}$ )

0 10 20 30 40 50 60

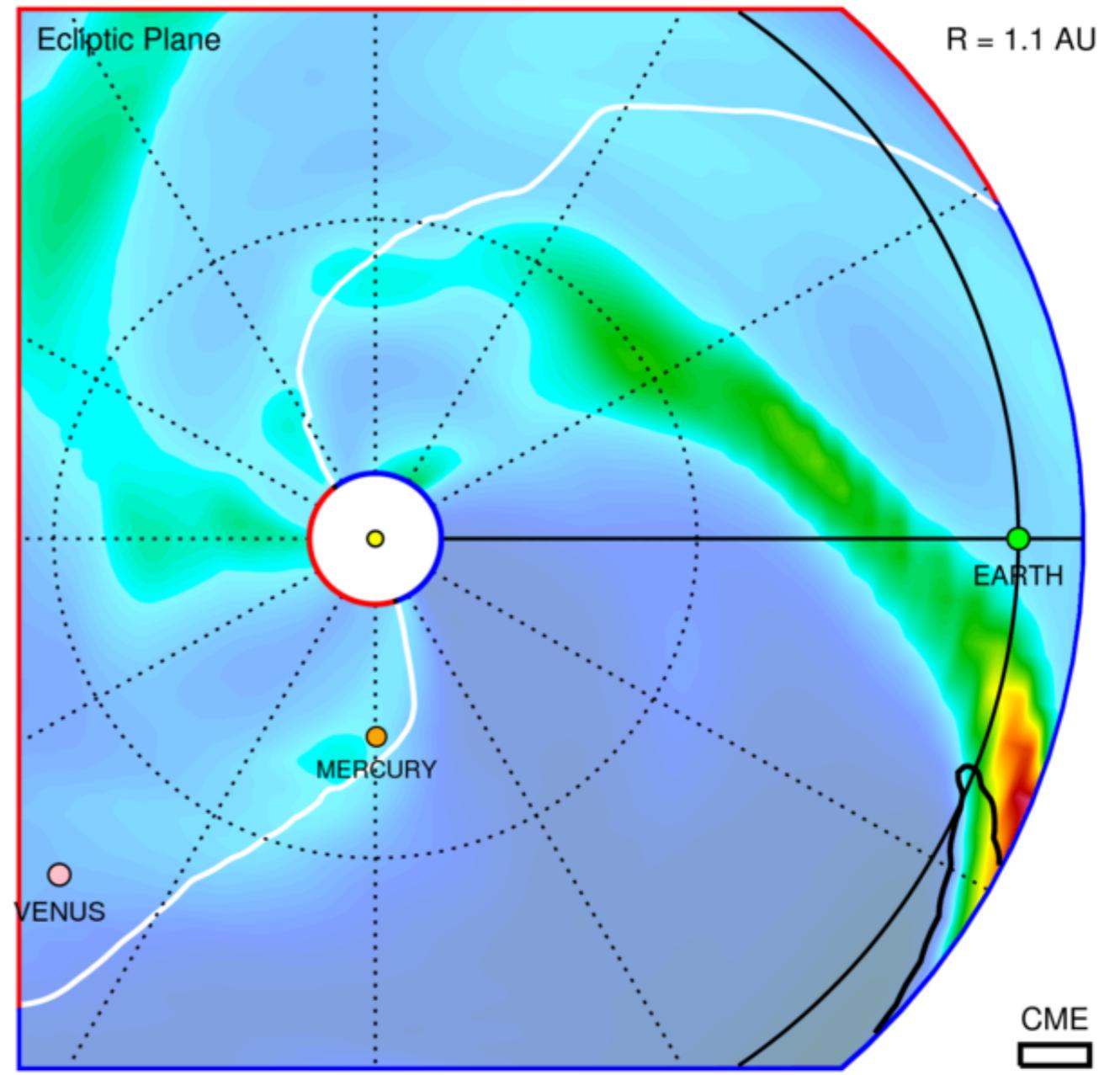
IMF polarity

- +

HCS

ENLIL-lowres + GONGb-WSAtu-Cone / a6b1-d4t1x1

2010-04-03T00



$r^2 N$  ( $\text{cm}^{-3}$ )

0 10 20 30 40 50 60

IMF polarity

- +

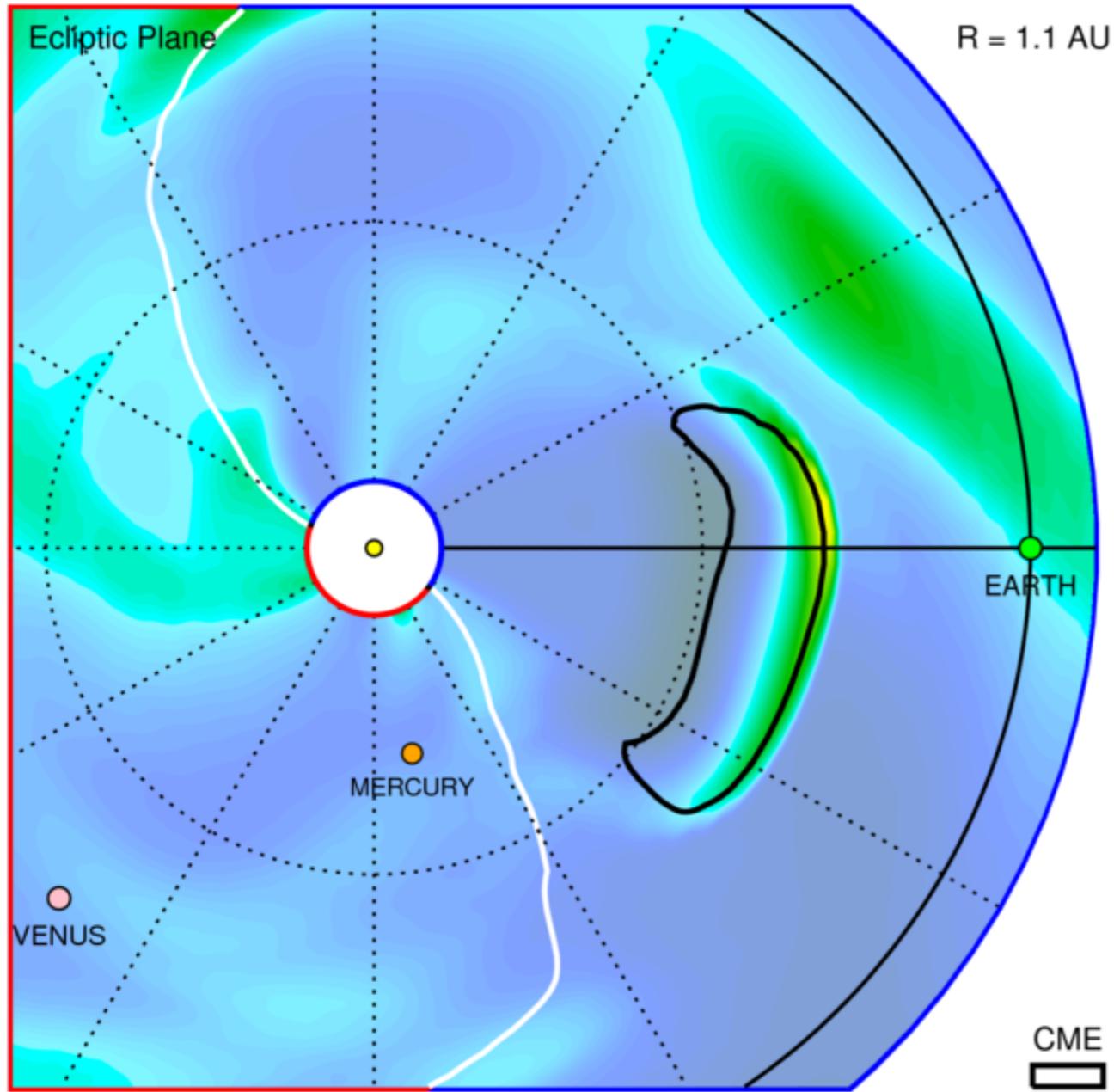
HCS

HELCASTS+DONKI : HelioWeather ENLIL-lowres + GONGz-WSAtu-Cone / a6b1-d4t1x1

HELCASTS+DONKI : HelioWeather

# WSA-ENLIL-DONKI-HELCCATS — Solar Wind Speed in 2012-03

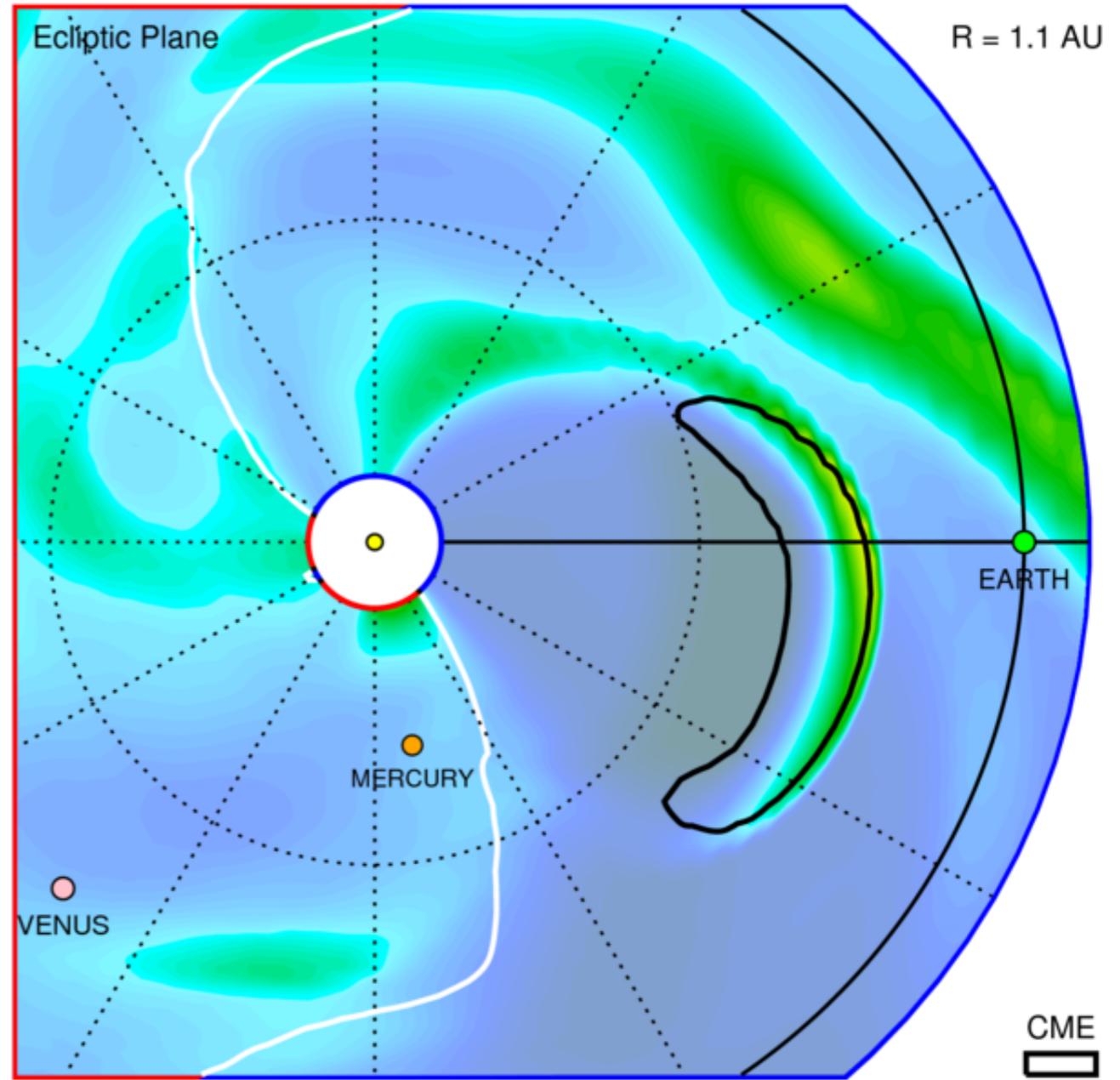
2010-04-05T00



$r^2 N$  ( $\text{cm}^{-3}$ ) 0 10 20 30 40 50 60

IMF polarity - + HCS

2010-04-05T00



$r^2 N$  ( $\text{cm}^{-3}$ ) 0 10 20 30 40 50 60

IMF polarity - + HCS

ENLIL-lowres + GONGb-WSAtu-Cone / a6b1-d4t1x1

HELCCATS+DONKI : HelioWeather ENLIL-lowres + GONGz-WSAtu-Cone / a6b1-d4t1x1

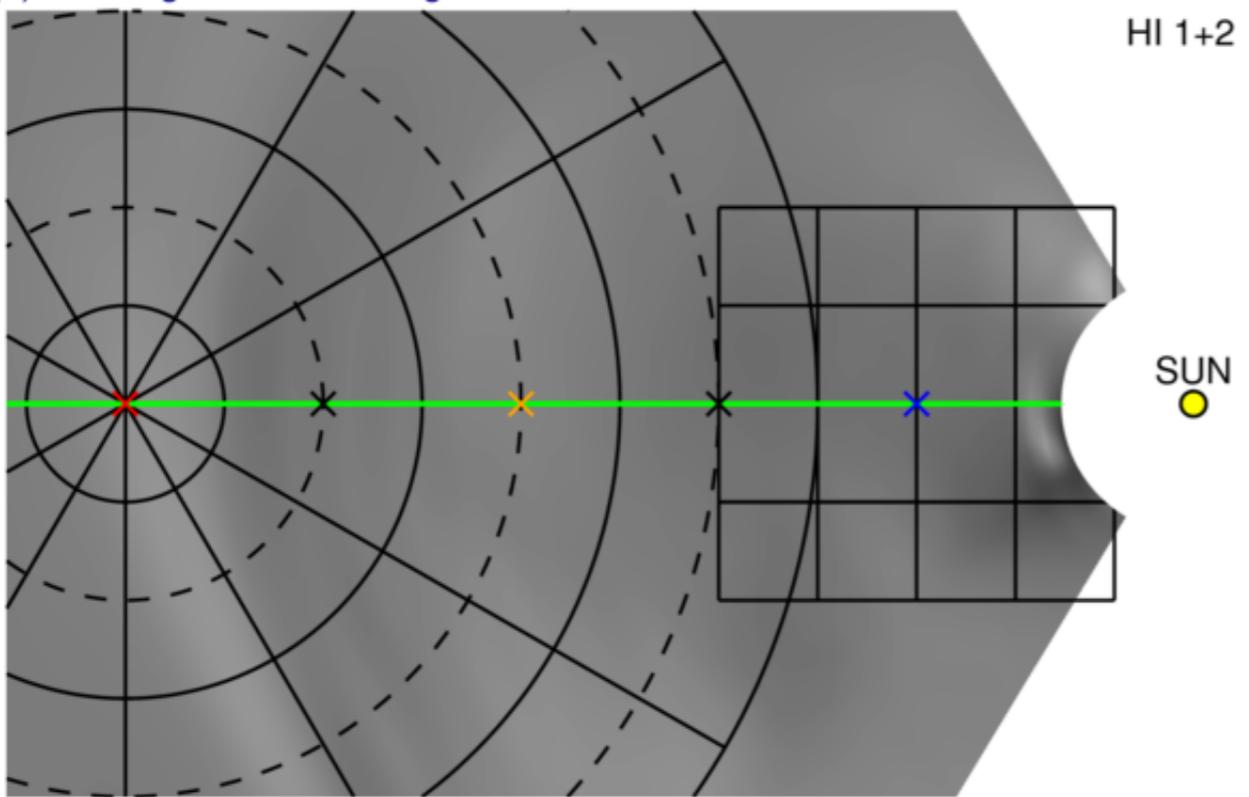
HELCCATS+DONKI : HelioWeather

# WSA-ENLIL-DONKI-HELCCATS — Solar Wind Speed in 2012-03

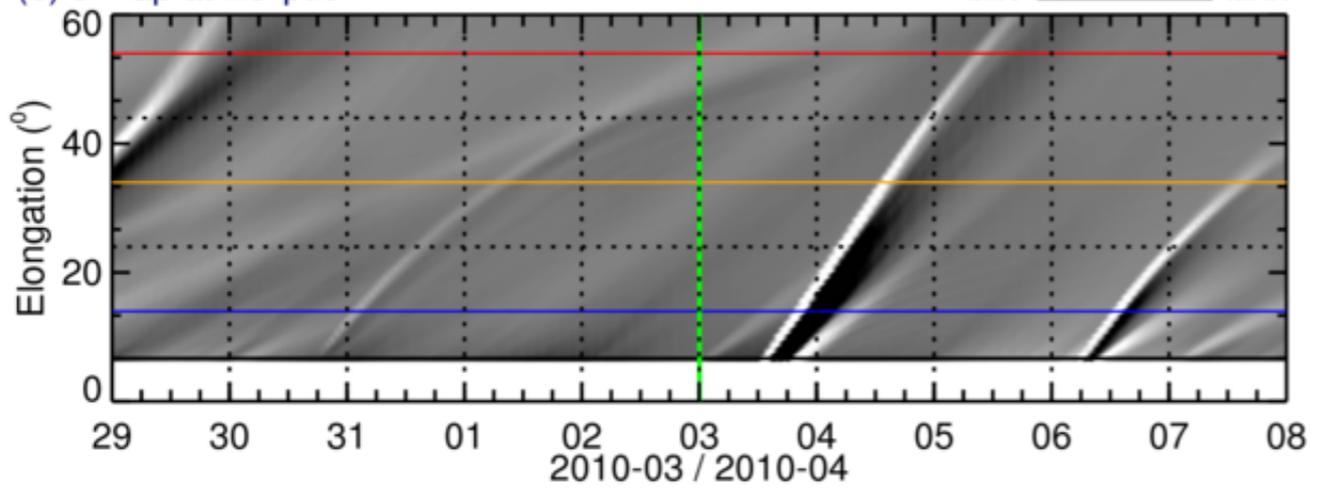
STEREO-A

2010-04-03T00:00

(a) Running-difference image



(b) J-map at Ecliptic



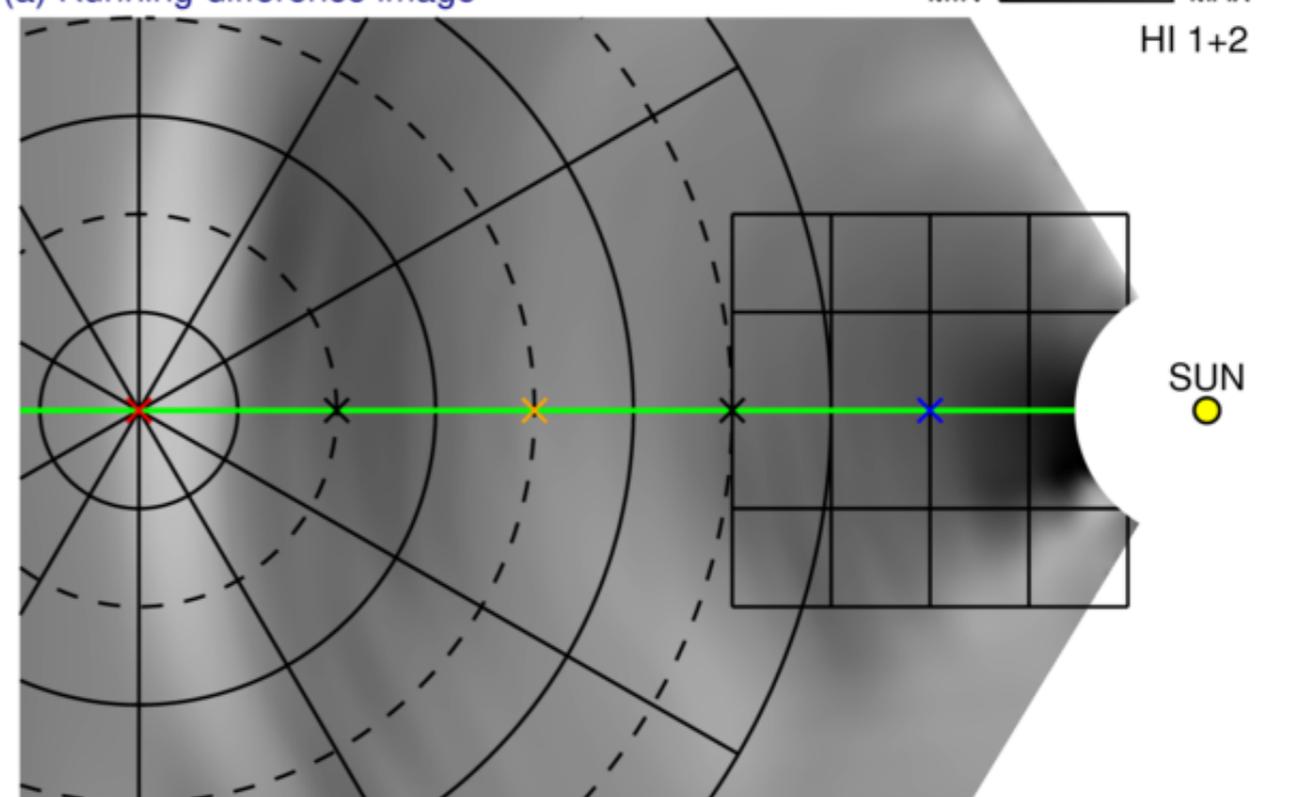
ENLIL-lowres + GONGb-WSAtu-Cone / a6b1-d4t1x1

HELCCATS+DONKI : HelioWeather

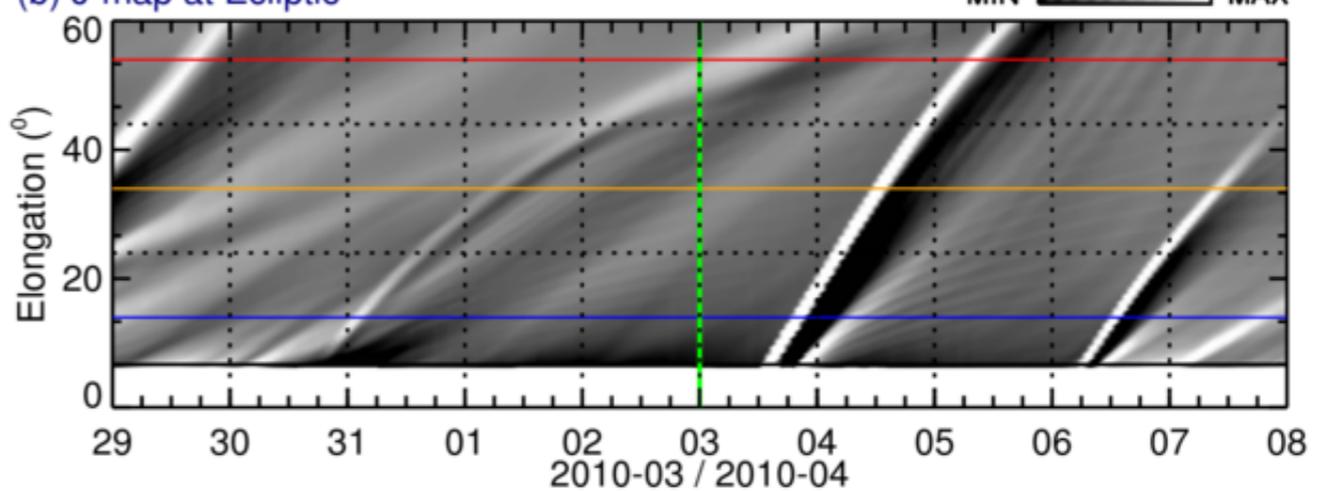
STEREO-A

2010-04-03T00:00

(a) Running-difference image



(b) J-map at Ecliptic



ENLIL-lowres + GONGz-WSAtu-Cone / a6b1-d4t1x1

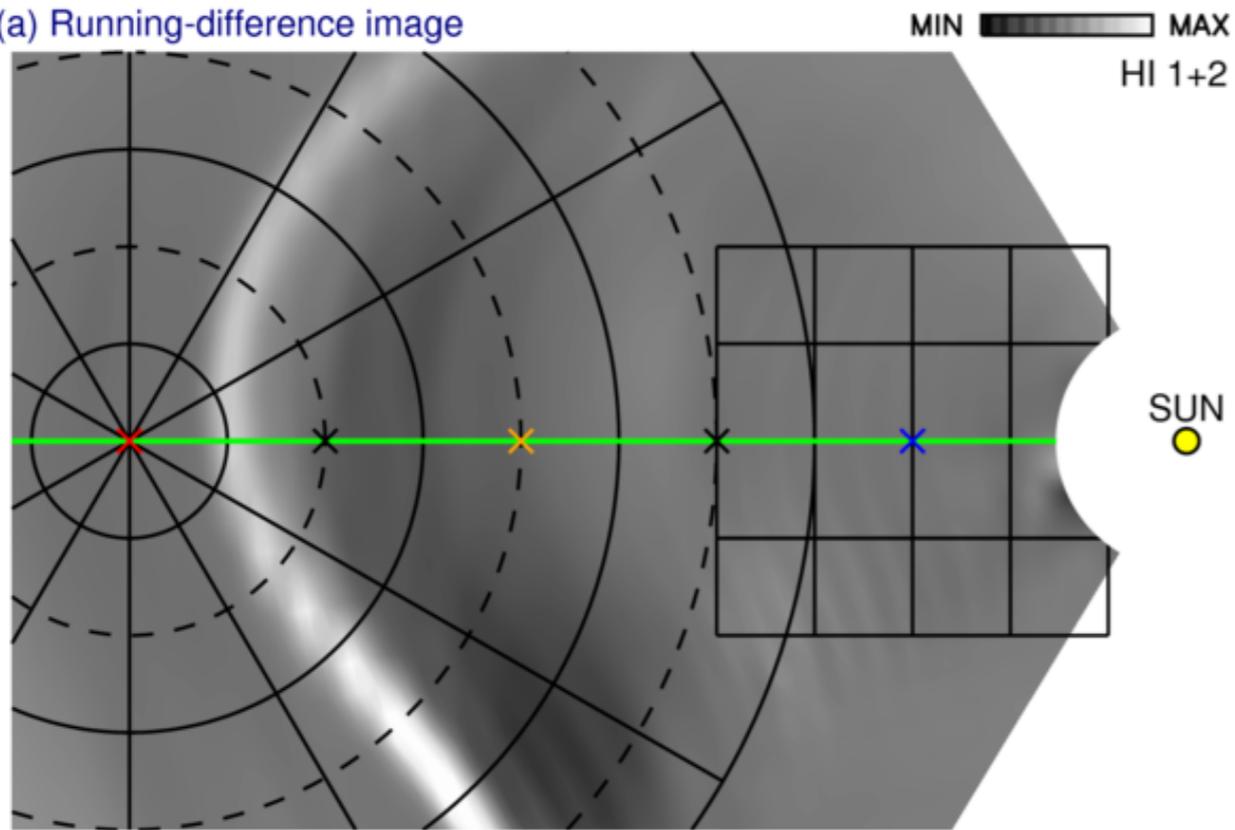
HELCCATS+DONKI : HelioWeather

# WSA-ENLIL-DONKI-HELCASTS — Solar Wind Speed in 2012-03

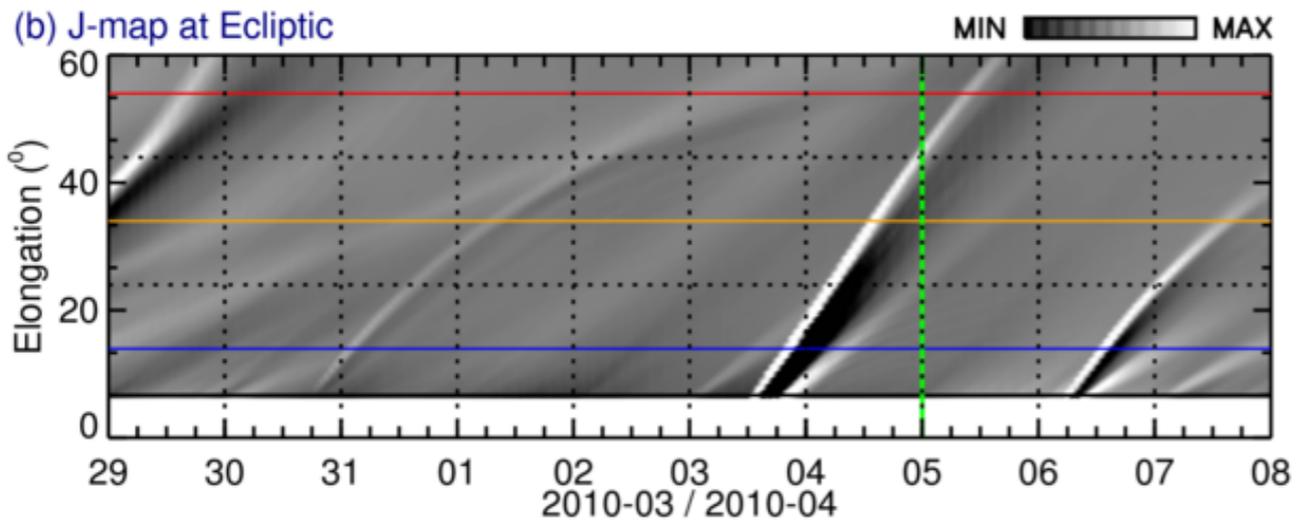
STEREO-A

2010-04-05T00:00

(a) Running-difference image



(b) J-map at Ecliptic



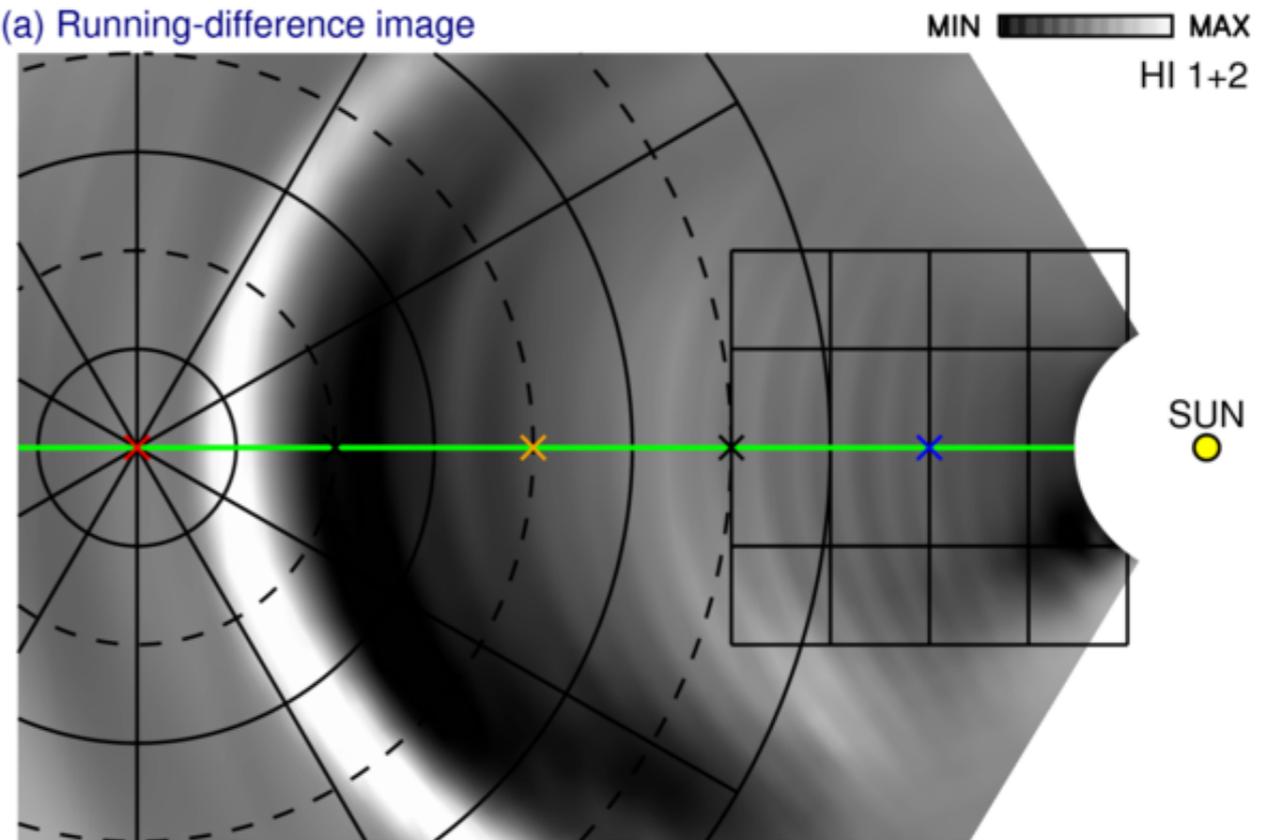
ENLIL-lowres + GONGb-WSAtu-Cone / a6b1-d4t1x1

HELCASTS+DONKI : HelioWeather

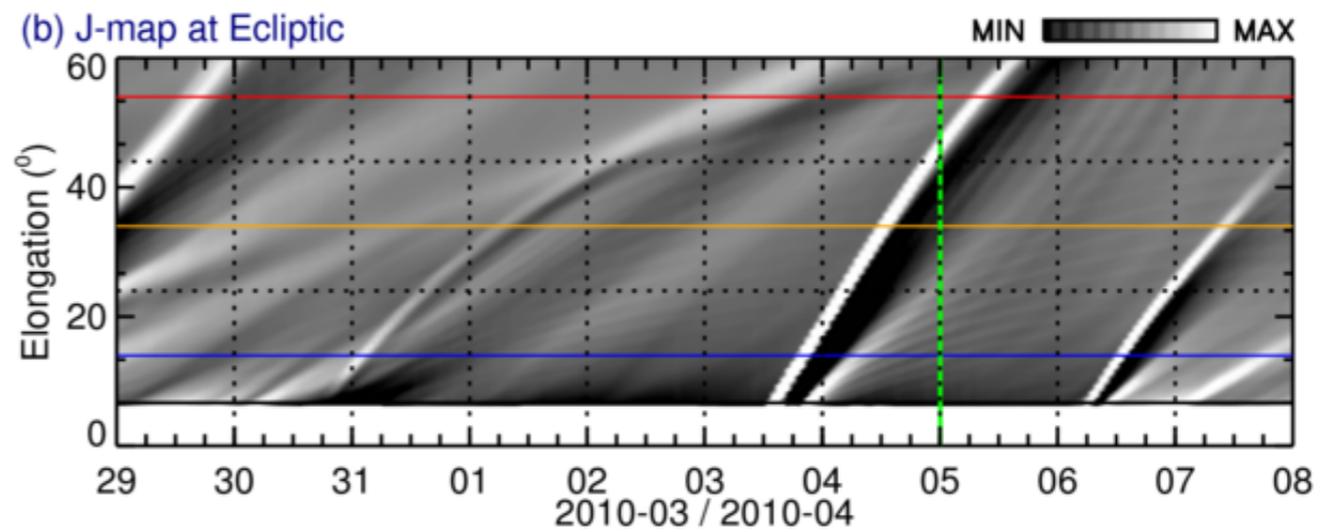
STEREO-A

2010-04-05T00:00

(a) Running-difference image



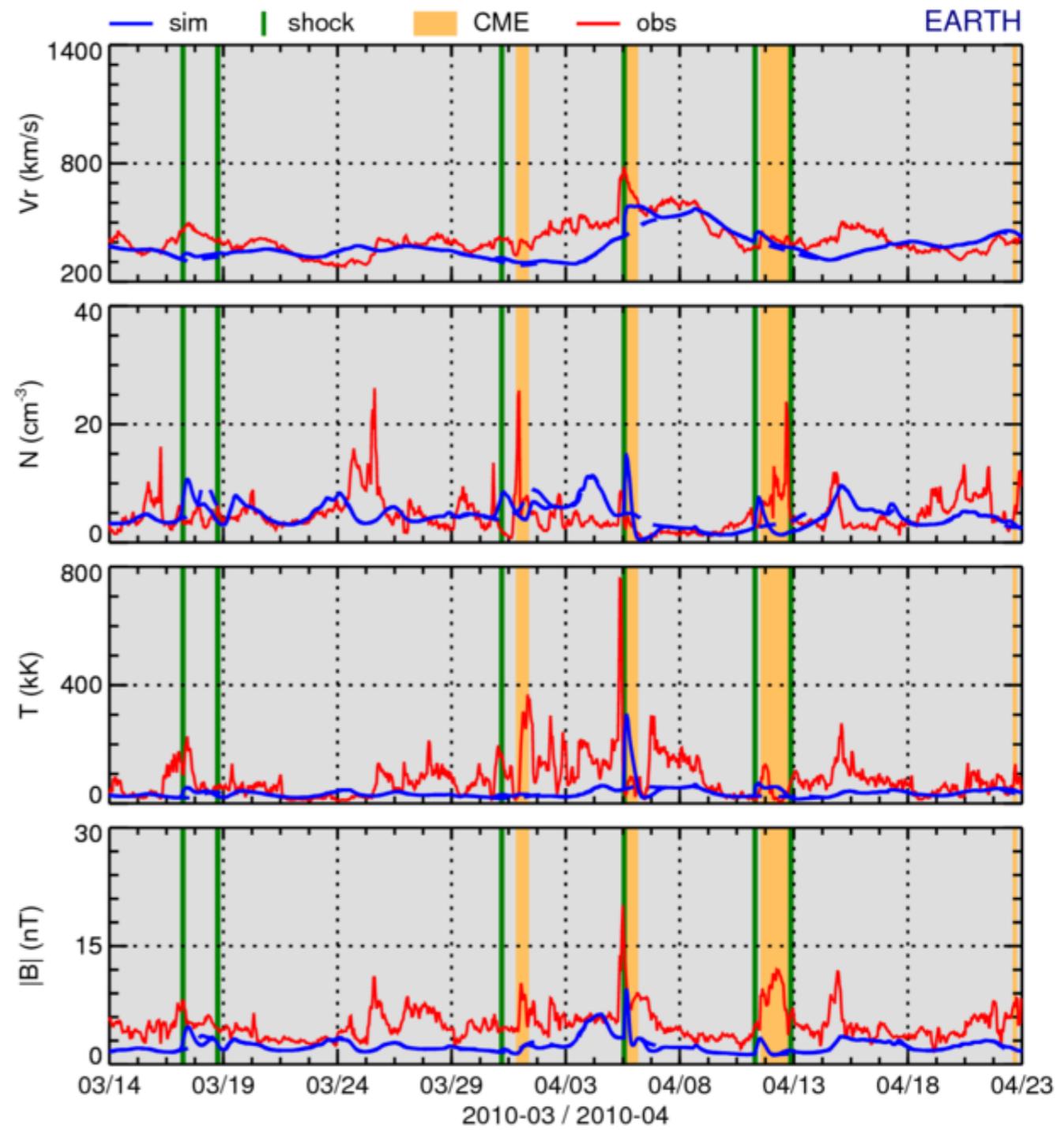
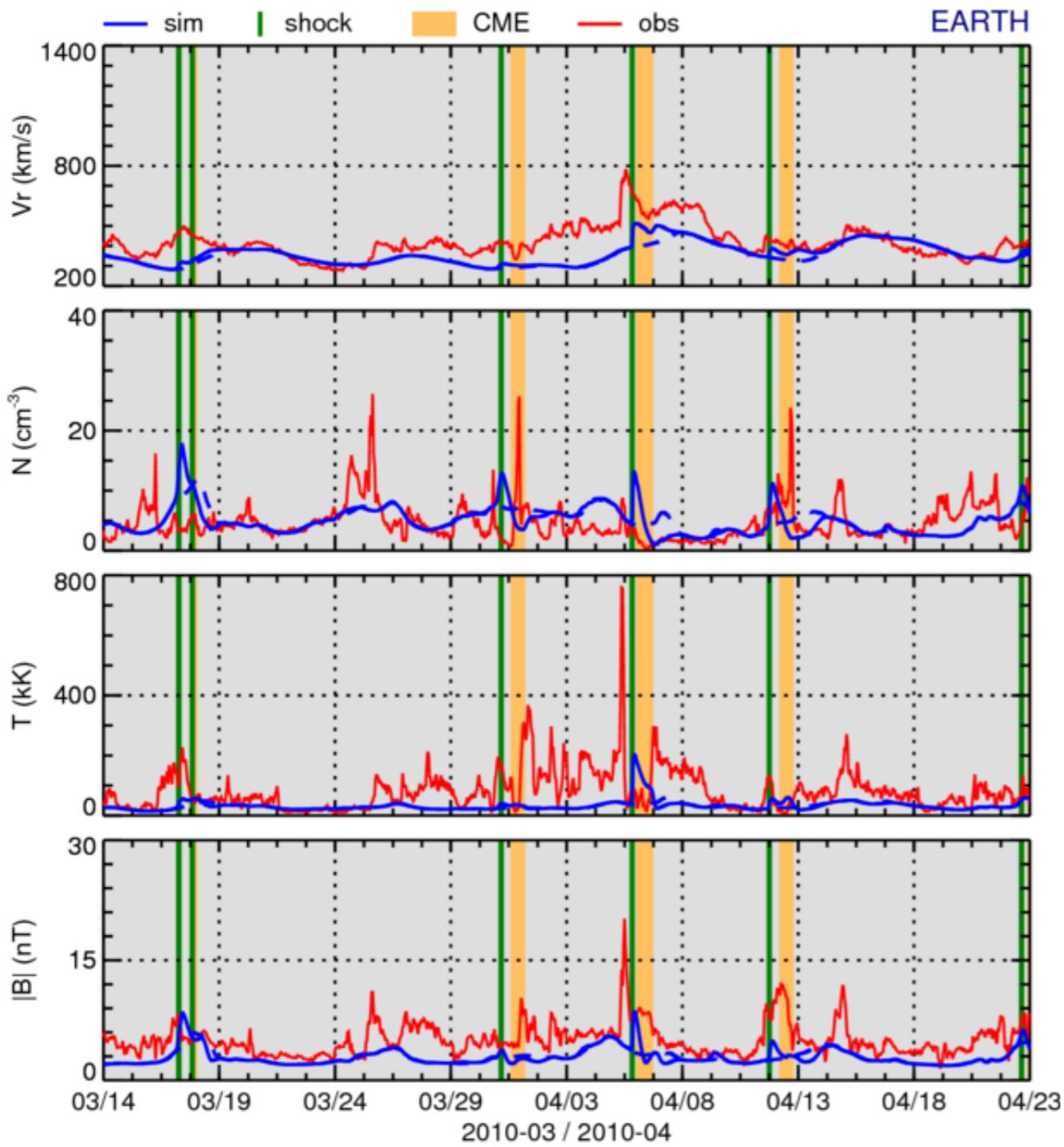
(b) J-map at Ecliptic



ENLIL-lowres + GONGz-WSAtu-Cone / a6b1-d4t1x1

HELCASTS+DONKI : HelioWeather

# WSA-ENLIL-DONKI-HELICATS — Solar Wind Speed in 2012-03



## WSAduz-SWRC – ENLIL Solar Wind Prediction – Velocity-2

Model Input:

[IPSbd](#)

[WSAdub-SWRC](#)

[WSAduz-SWRC](#)

[WSAdtb-SWRC](#)

[WSAdwb-SWRC](#)

Model Output:

[Velocity-1](#)

[Velocity-2](#)

[Density-1](#)

[Density-2](#)

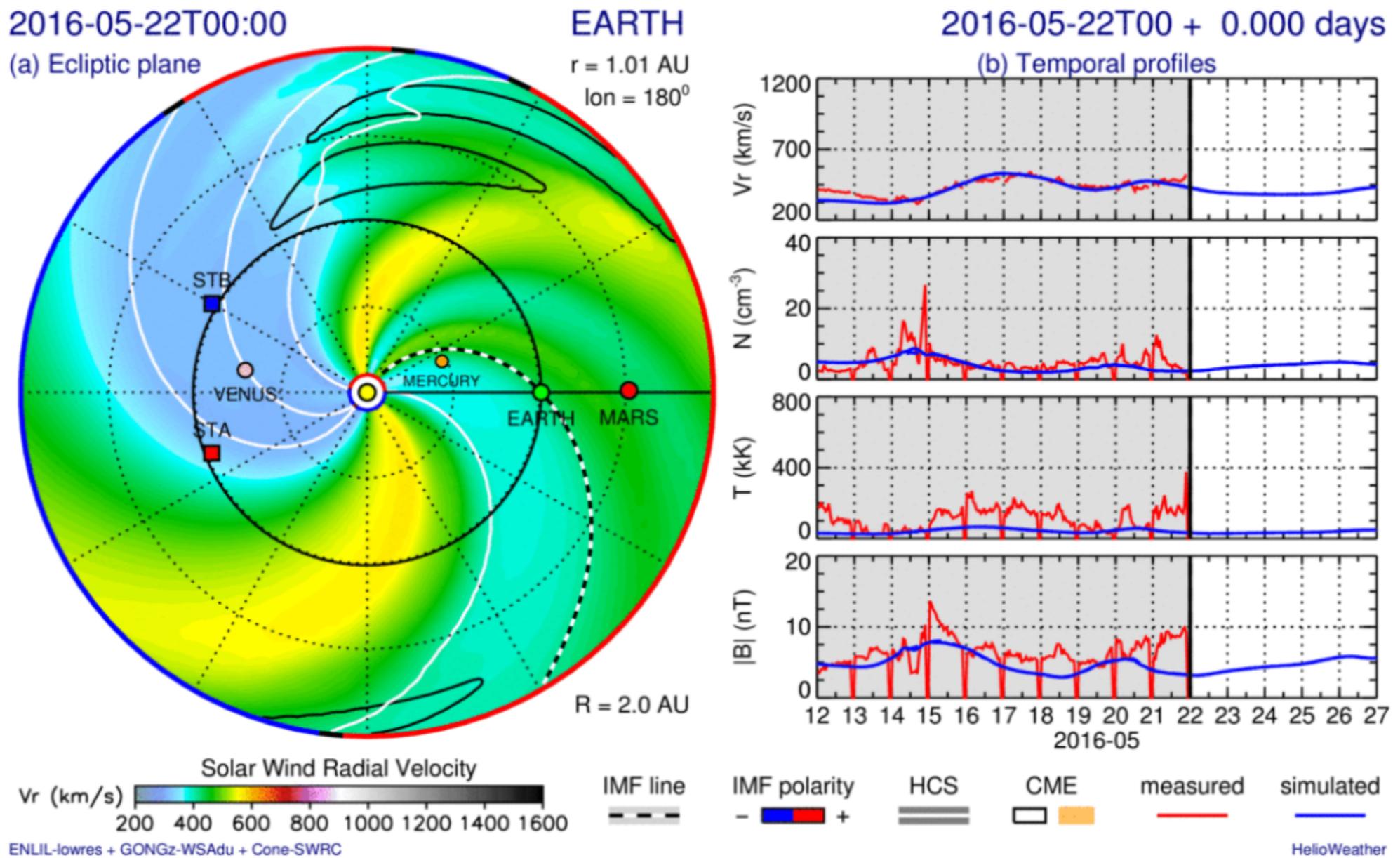
[STEREO-A](#)

[STEREO-B](#)

[Shock-SEP](#)

[Evolution](#)

[J-Maps](#)



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[INFO](#)



+00T00 ▾

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- Various models & various outputs — continuous near-real time verification
- Fixed model parameters do not reflect variations over the solar cycle

# International Use of ENLIL for Operational Space Services

*(input T. Onsager, NOAA/SWPC)*

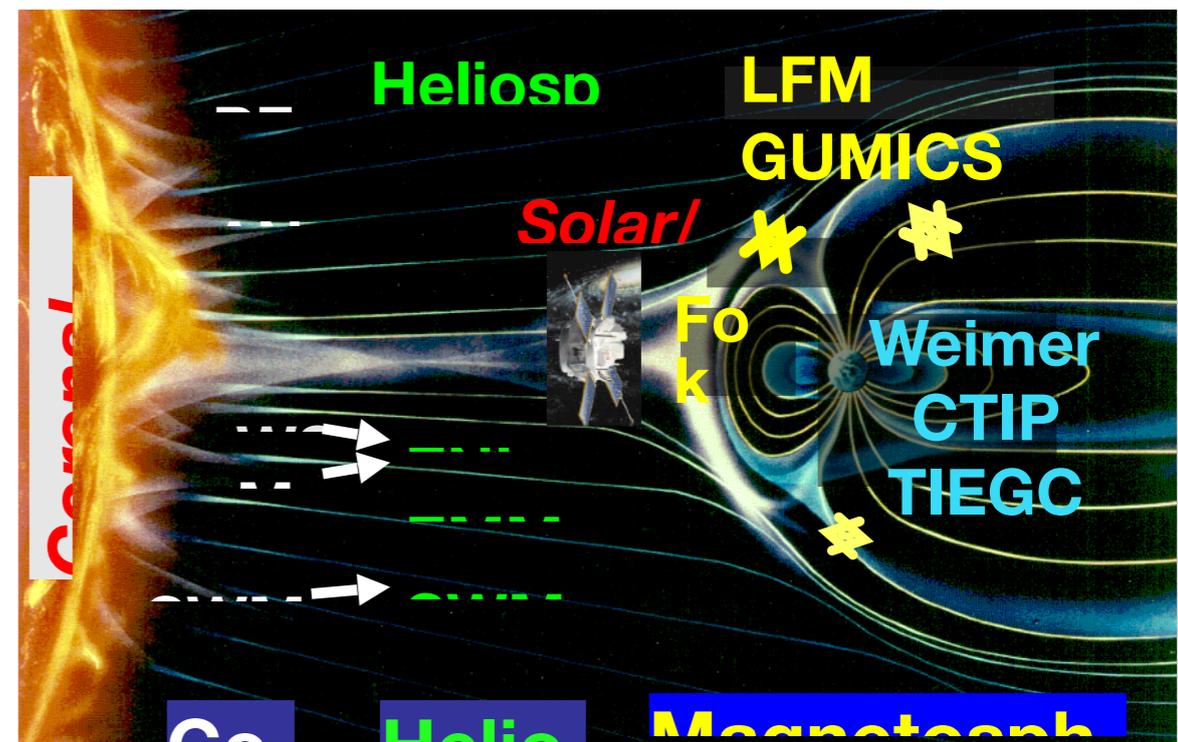
- Enlil is utilized internationally for operational prediction of the background solar wind and CME propagation and is an essential forecasting tool
- Memorandums of Understanding and non-disclosure agreements with operational partners have been signed with Enlil as an element of cooperation
- Software versioning is maintained to ensure consistent product
- In operations, multiple runs may be evaluated for official forecasts
- Current partners include:
  - United Kingdom (Met Office)
  - Korea (RRA/KSWC)
  - Australia (BoM/SWS)
- Additional international partners have expressed interest



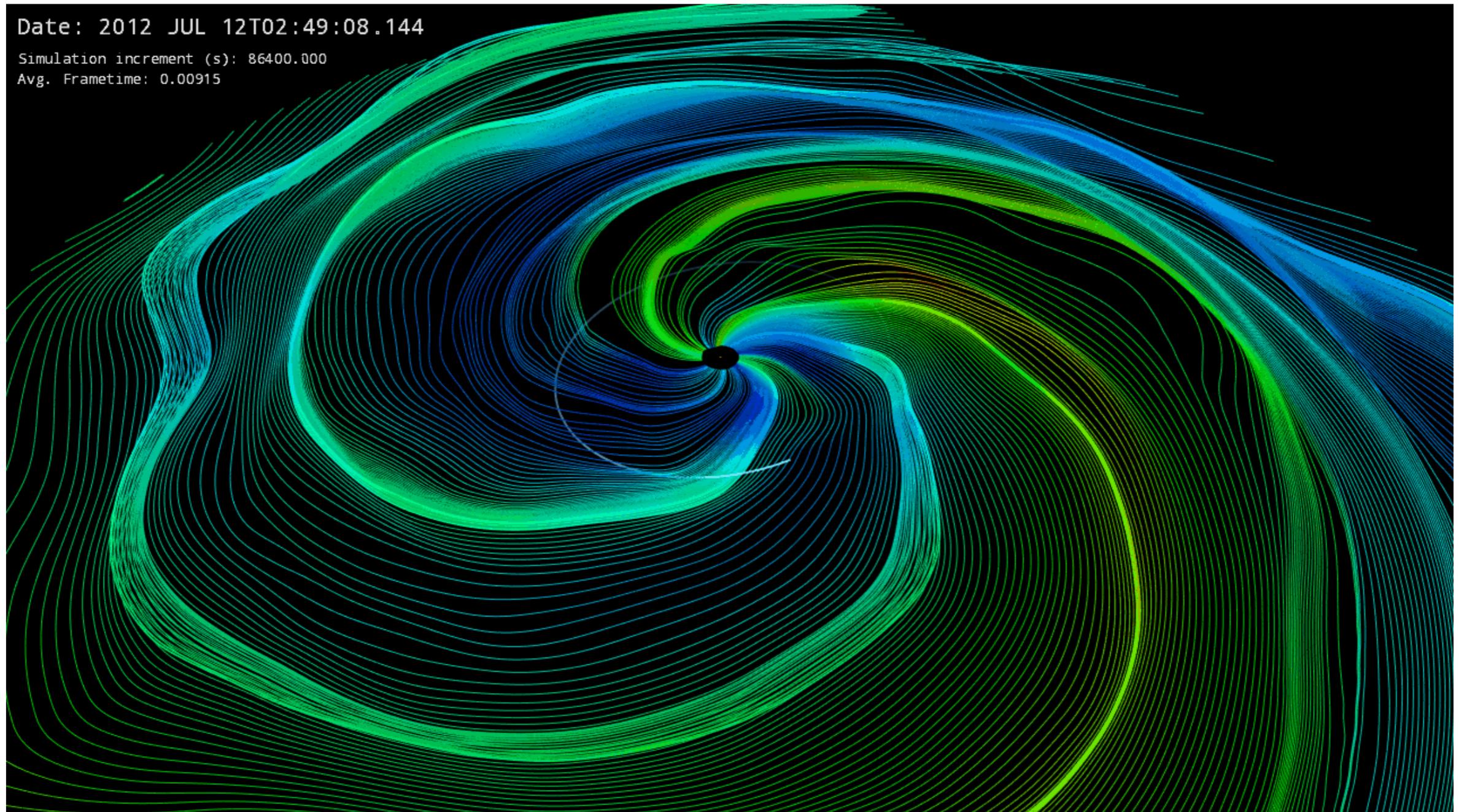
# ENLIL at NASA/CCMC

*(input from Anna Chulaki and L. M. Mays, NASA/CCMC)*

- Community Coordinated Modeling Center (CCMC) is a multi-agency facility operated at NASA/GSFC
- ENLIL was implemented at CCMC in 2005 (legacy of NSF/CISM project) and it is used continuously since then for run-on-request & mission support services
- Total number of ENLIL requests is 5703 (as of May 17, 2016)
- There are only 9 papers & 22 presentations in CCMC database that explicitly mention ENLIL by name in their title. This is probably a very small subset of papers/publications that utilized ENLIL
- CCMC staff also independently evaluates and validates the modeling system and uses it for space weather predictions to support NASA robotic missions
- It is also used for education and public outreach



# Visualization of Heliospheric Simulations at Planetariums



- OpenSpace is a multi-scale rendering software platform
- OpenSpace is open source software available for MacOD, PC, Linux on laptops, desktops, classrooms, planetariums, virtual reality
- OpenSpace enables interactive 3D visualization of space weather models at the NASA/CCMC
- ENLIL will be featured in the Sun-Earth day planetarium at AMNH (New York) on July 27, 2017

# Visualization of Heliospheric Simulations at Planetariums



Morrison Planetarium (San Francisco, CA)



Hayden Planetarium (New York, NY)

- Collaborative project between NASA/CCMC, Linkoping Univ (Sweden), and American Museum of Natural History (New York, NY)
- OpenSpace tool — interactive, multi-scale, multi-display environments to explore our current understanding of Universe
- Ultimate goal — a general open-source visualization software enabling synchronized presentations in planetariums across the globe

# Summary

- We developed new versions of the WSA-ENLIL-Cone modeling system capable to handle new inputs & produce new outputs
- We set up the project testbed system for “on-fly” verifications of various new versions under near-real-time conditions
- Updated WSA-ENLIL-Cone modeling system can now routinely predict:
  - ICME arrival times (ejecta and/or shock) in mid-heliosphere
  - ensemble modeling
  - evolving background solar wind
  - IMF topology and shock parameters for SEP models & alert plots
  - synthetic white-light images (for “mid-course” correction)
- We realized large-scale calibration & validation studies to evaluate new features, compare with previous versions, and with other models implemented at CCMC
- We continued in development of the SEP model and applied it to selected multi-CME event scenarios observed during maximum of solar activity
- We authored or co-authored 18 papers and made numerous presentations at domestic and international conferences. Impact of our work has been enhanced by CCMC users.
- Updated WSA-ENLILCone modeling system facilitates:
  - direct comparison with remote and in-situ observations at planets and spacecraft
  - high-quality images and animations red for presentations/publications
  - scripting system to support research and prediction activities
  - supports heliospheric predictions & mission planning relevant to NASA missions